

Numerical simulations

in Astronomy based on the

Particle Hydrodynamics

Seung-Hoon Cha
&
Matt A. Wood

Department of Physics and Astronomy
Texas A&M University-Commerce

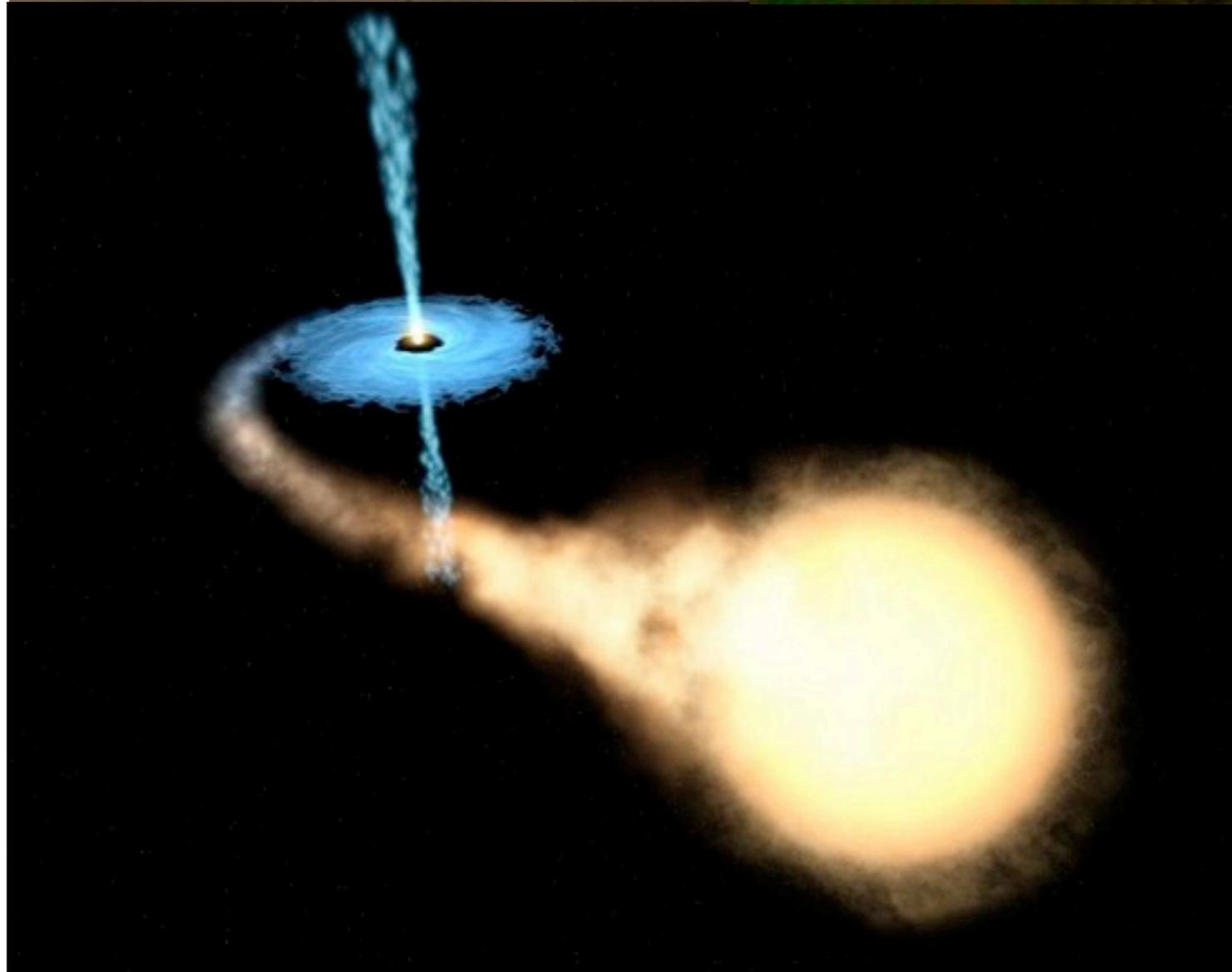
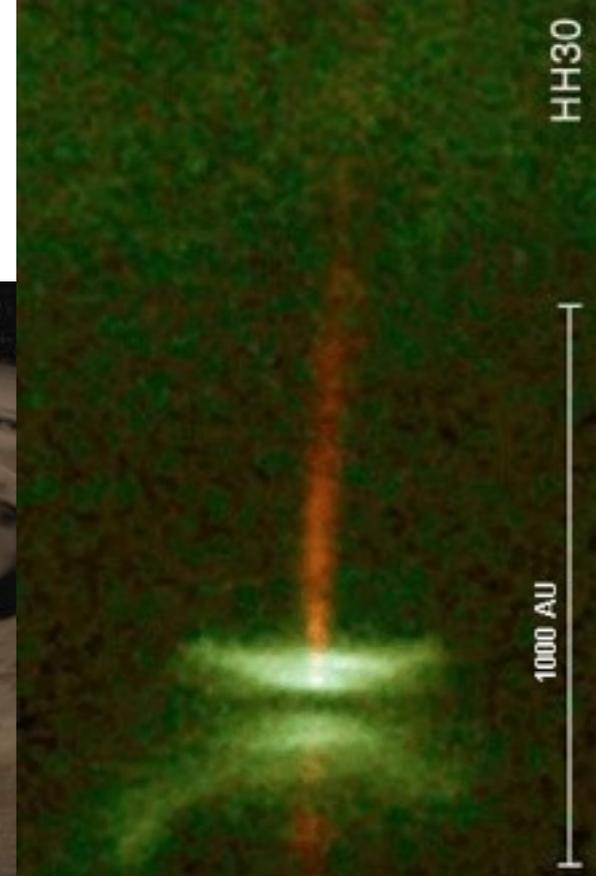
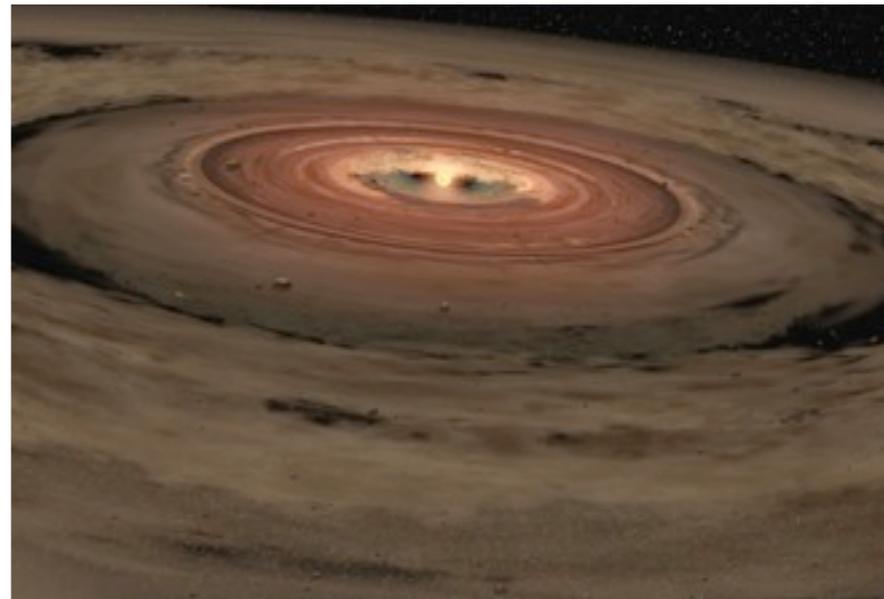
What we solve...

$$\frac{Dq}{Dt} = -V \nabla S,$$

$$q = \begin{cases} V \\ \mathbf{v} \\ e \end{cases}, \quad S = \begin{cases} \mathbf{v} \\ P \\ P\mathbf{v} \end{cases}$$

$$\nabla^2 \Phi = 4\pi G \rho \text{ and EOS.}$$

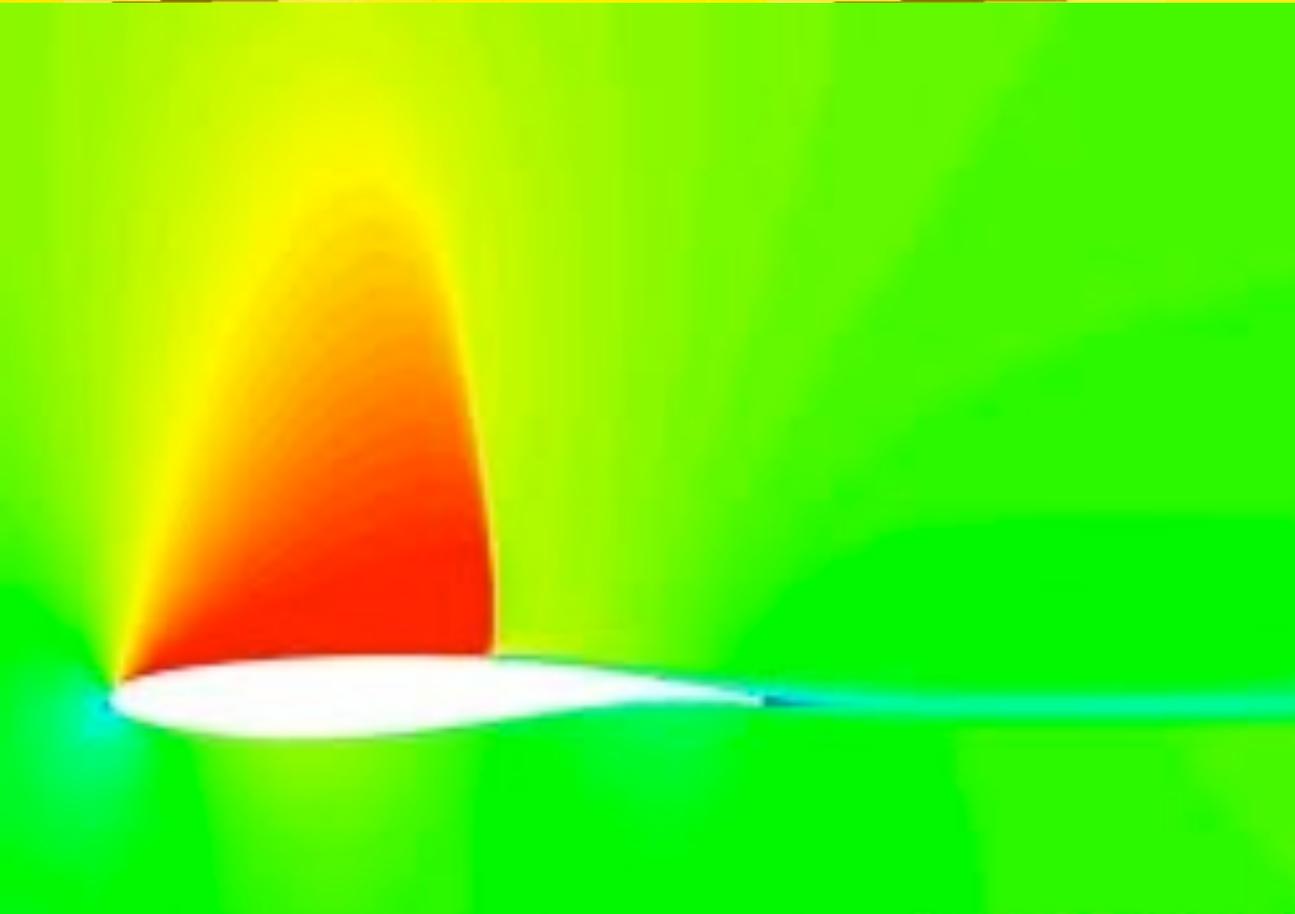
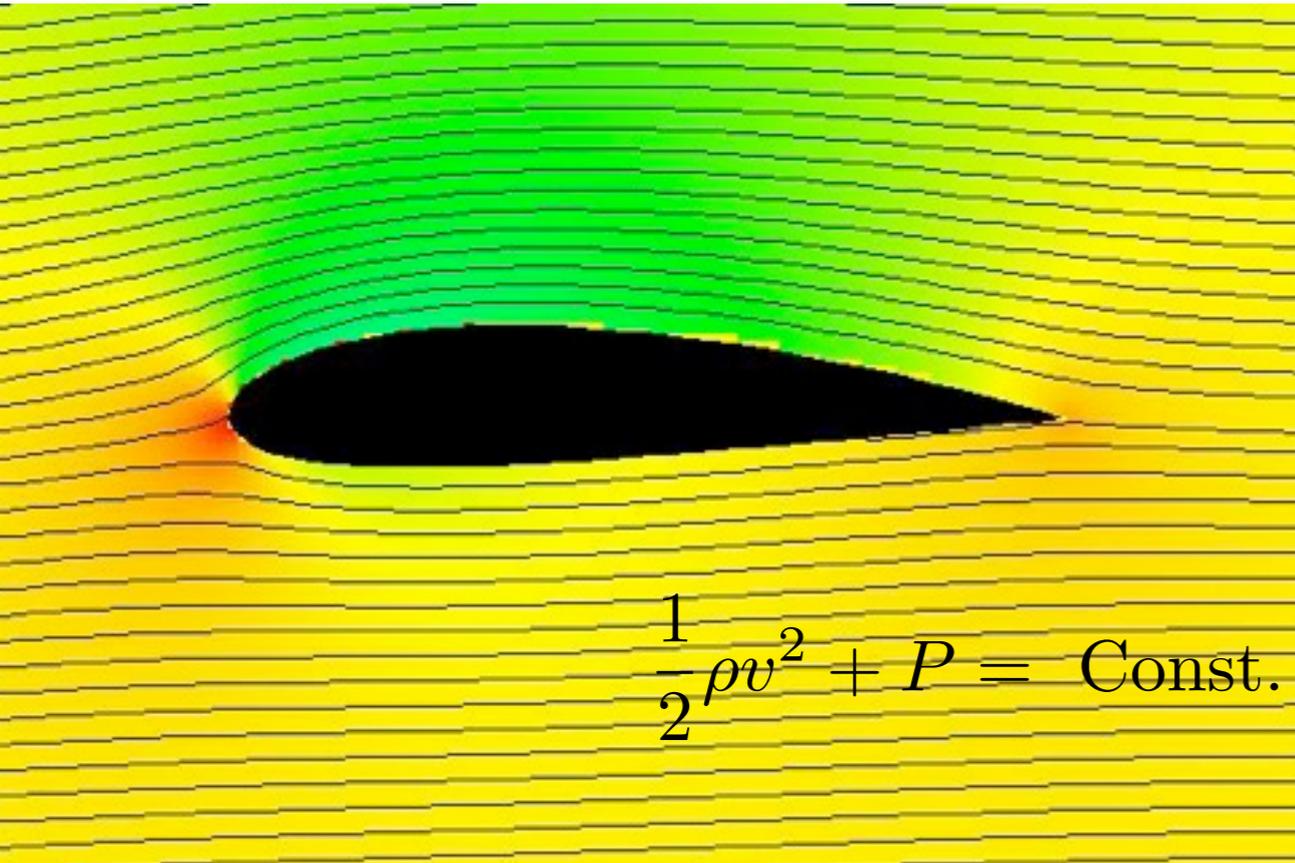
No general solution !



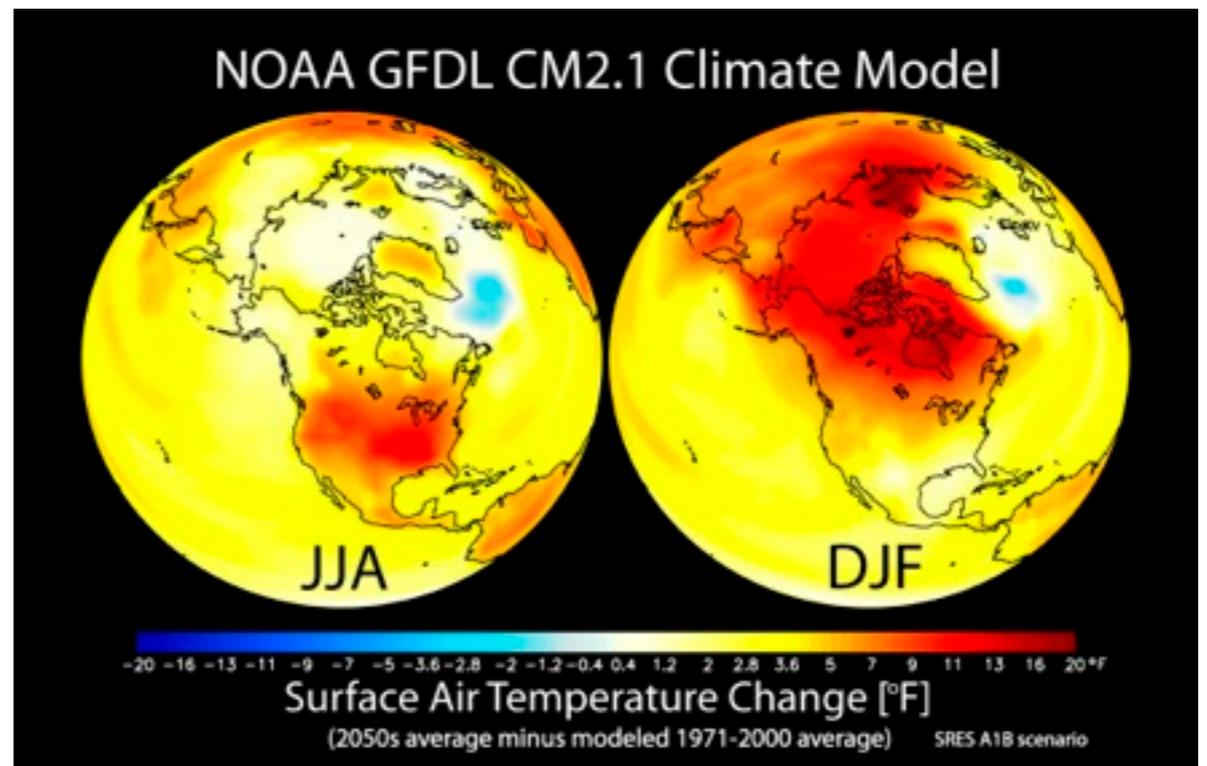
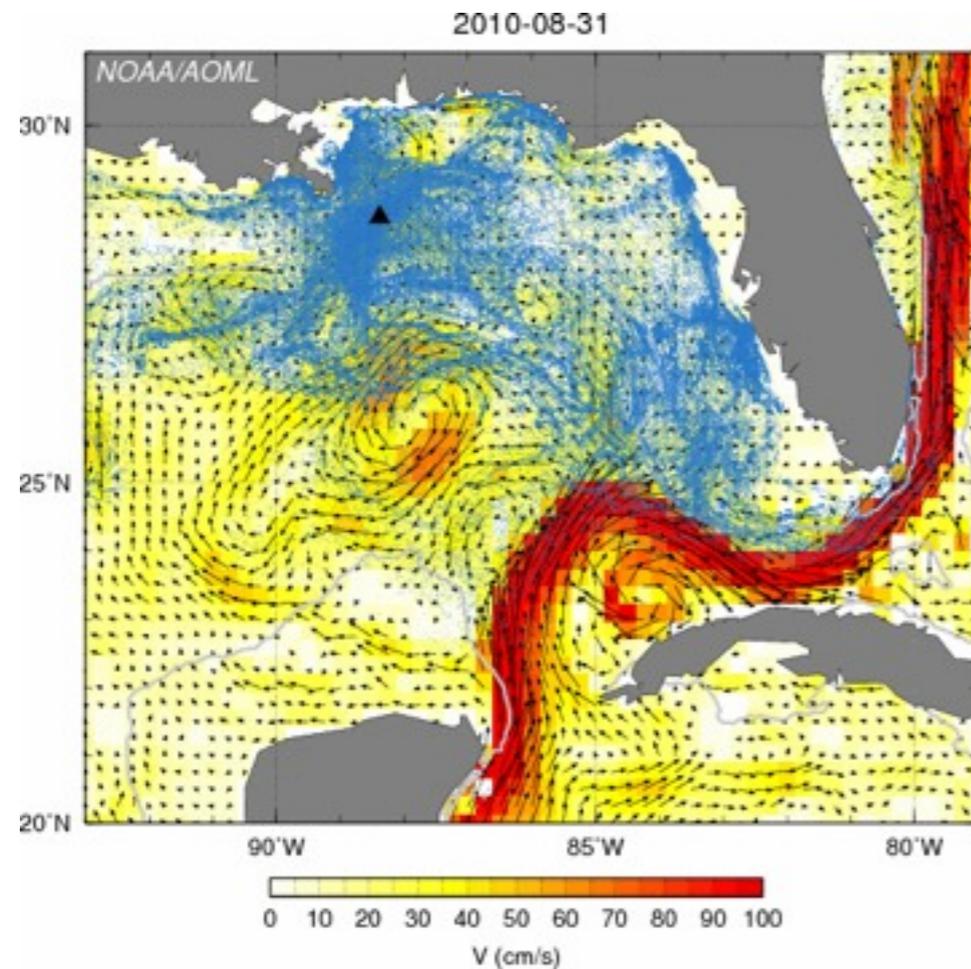


Transonic point

Bell X-1 exceeded
800 mph, M1.06 (1947)



All sciences use numerical simulations



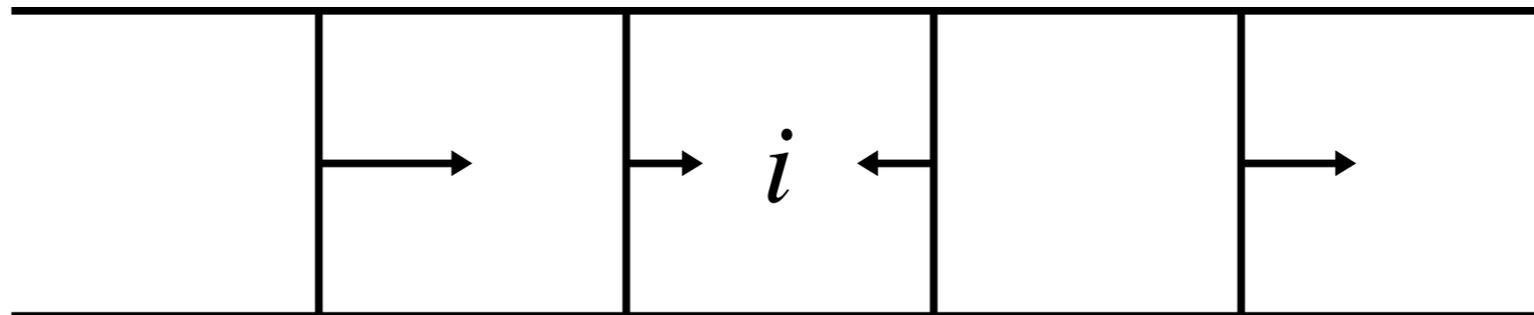
Astronomy is special **in the strong shocks.**

Why are the results different ?

Can I believe my simulation ?

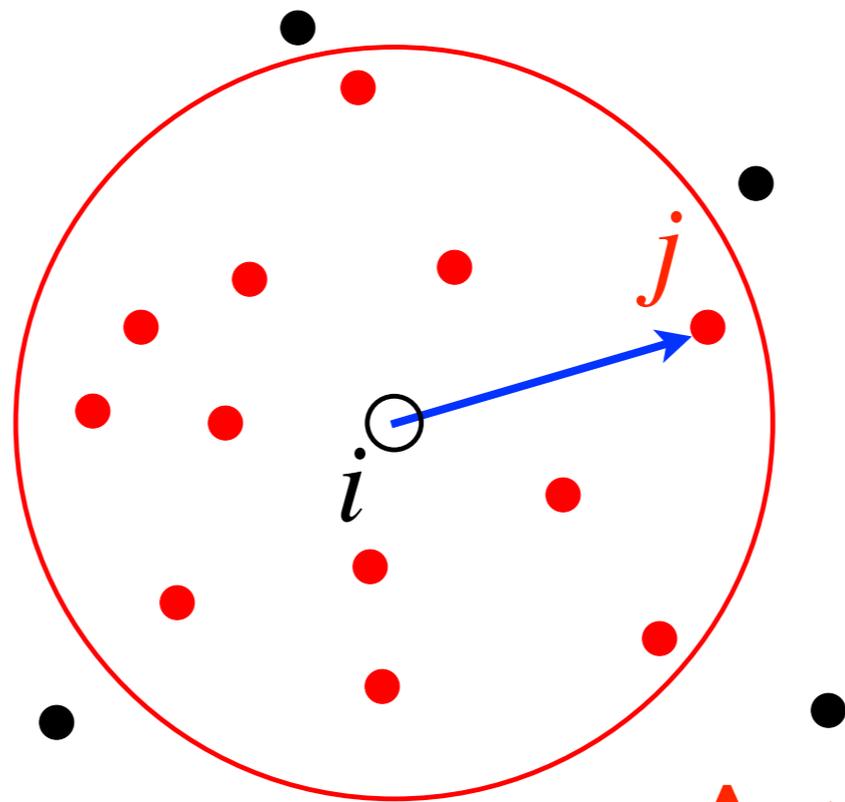
Finite Difference Method (FDM)

Grid-base code



$$\frac{\partial q_i}{\partial t} + \nabla F_i = 0$$

Smoothed Particle Hydrodynamics (SPH)



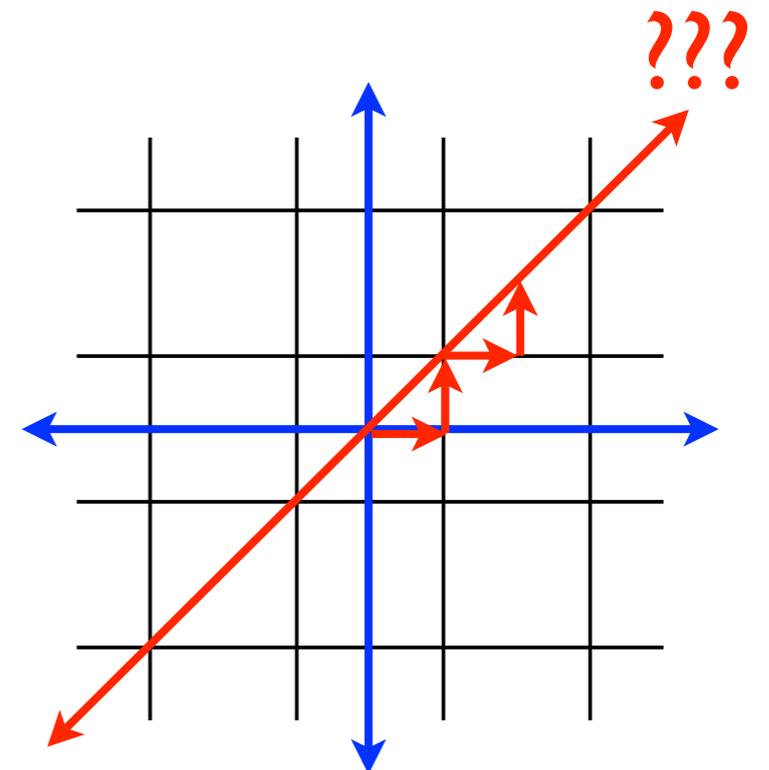
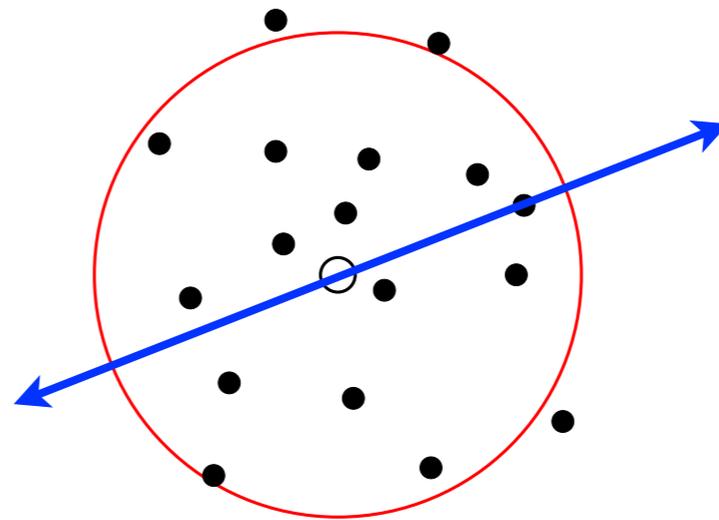
Artificial viscosity

$$\dot{\mathbf{v}}_i = - \sum_j m_j \left(\frac{P_i}{\rho_i} + \frac{P_j}{\rho_j} + \Pi_{ij} \right) \nabla_i W(\mathbf{r}_i - \mathbf{r}_j, h_{ij})$$

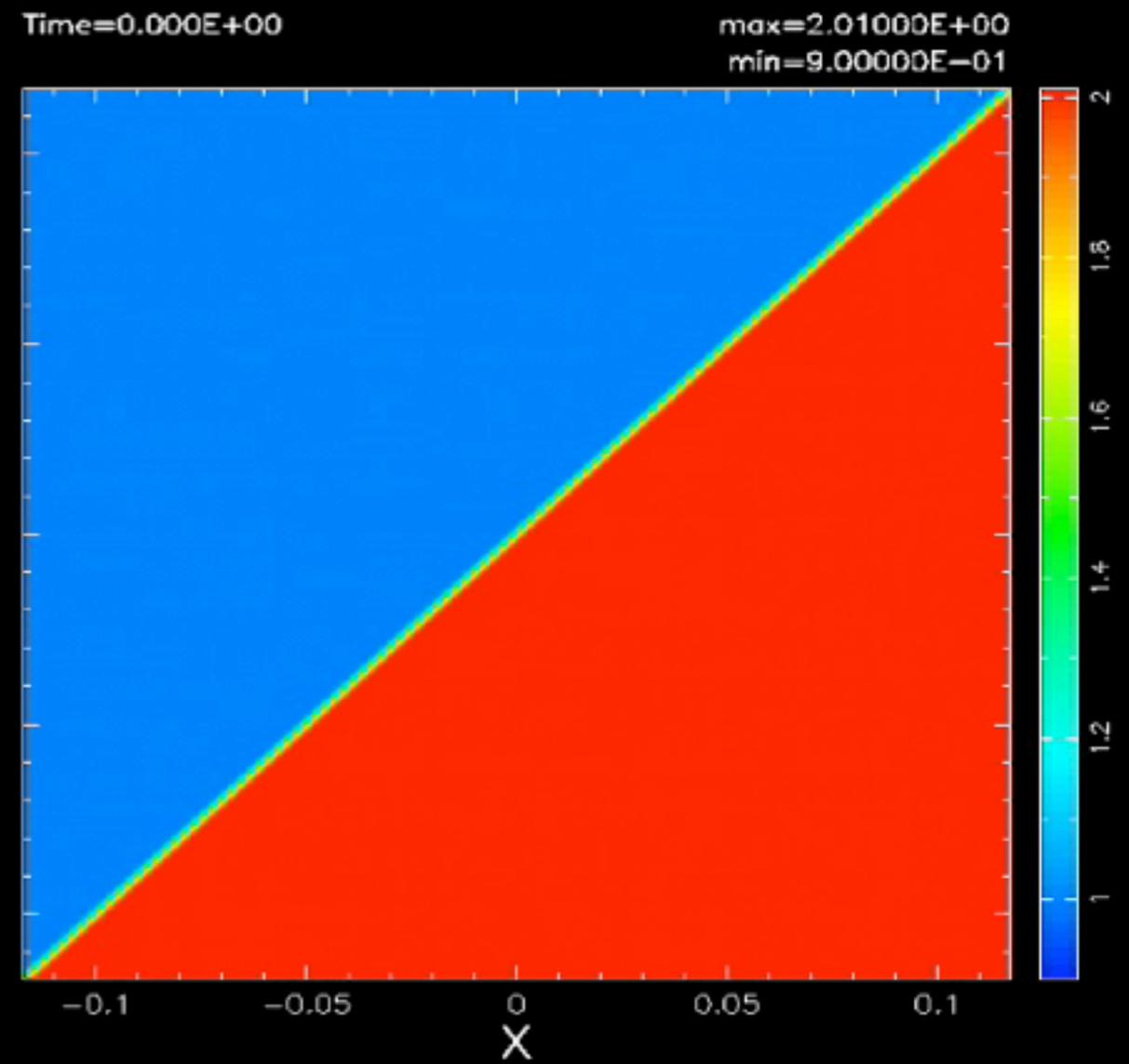
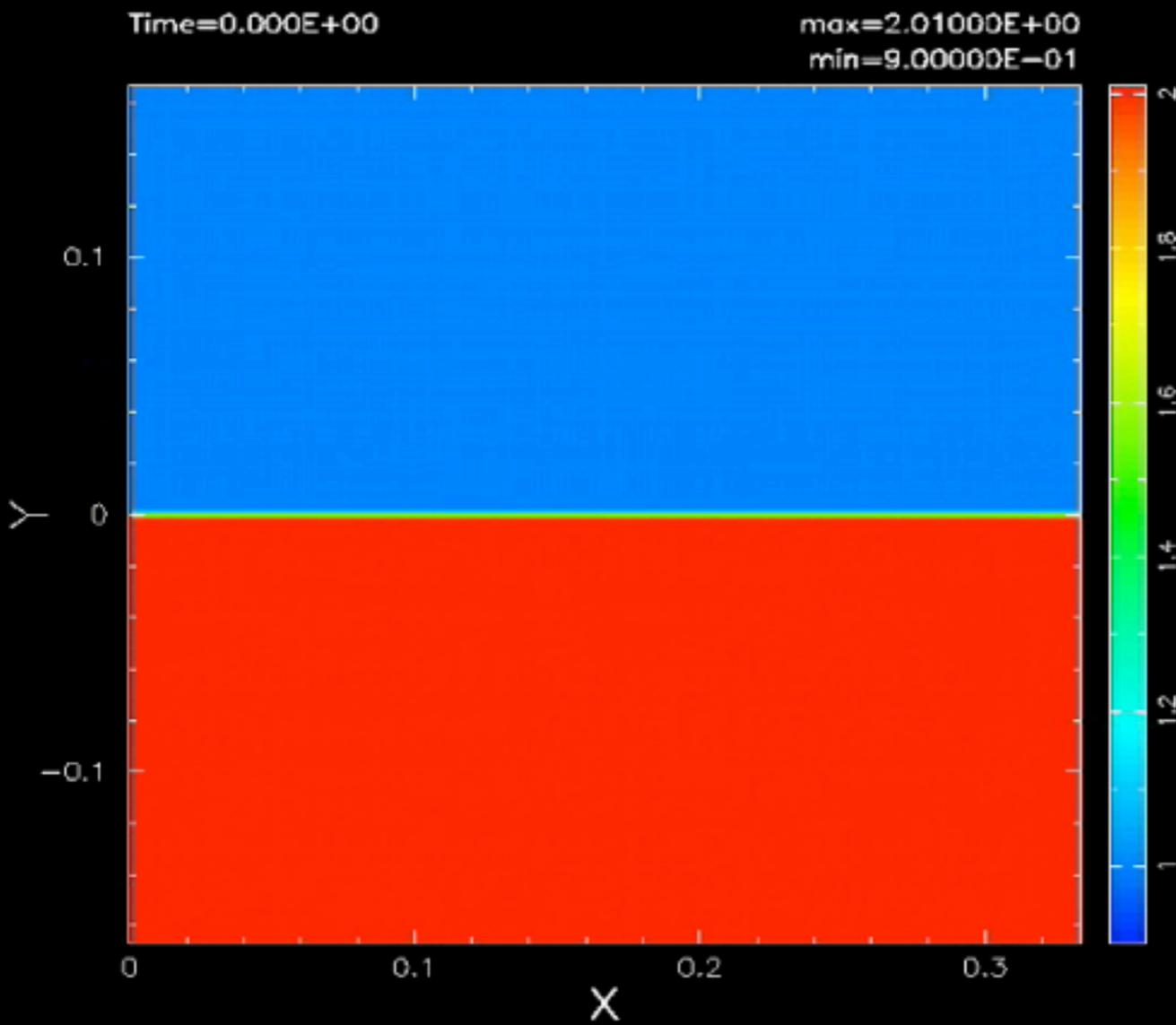
Why particle code ?

I. Multi-dimension

- 1-Dim. even in 3-Dim.
- Lagrangian .vs. Eulerian



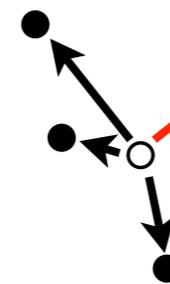
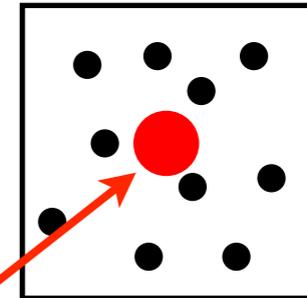
Kelvin-Helmholtz Inst.



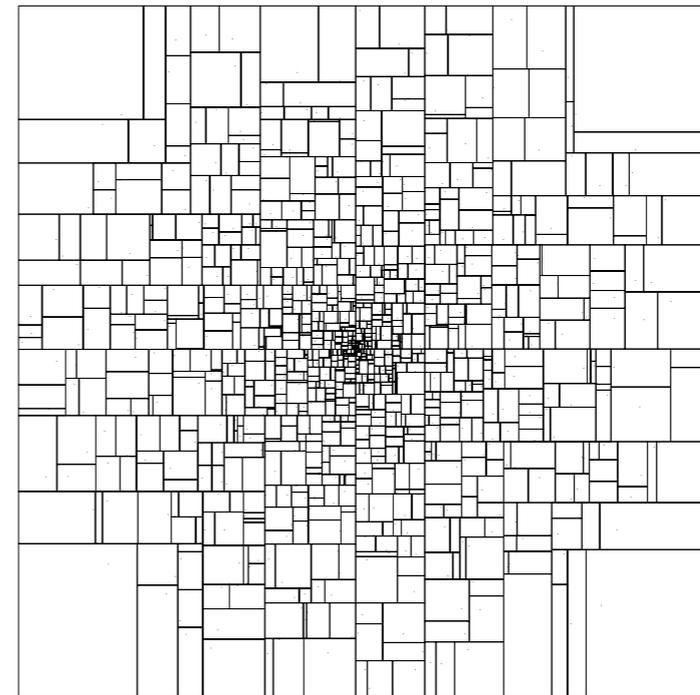
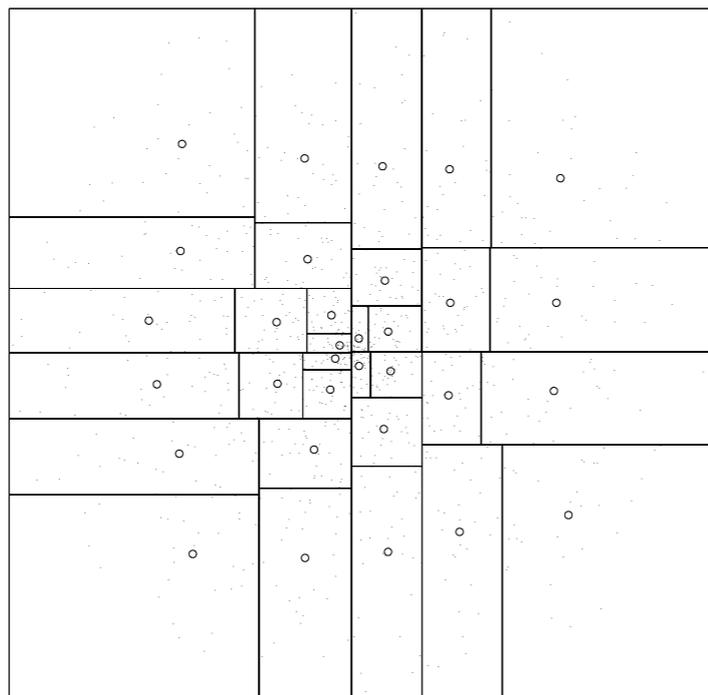
Why particle code ?

II. Self-gravity

- Accurate in neighbors, and fast in remote
- Binary tree : Recursive Bisection Method (in parallelization)



$N \log N$ instead of N^2



FDM vs SPH

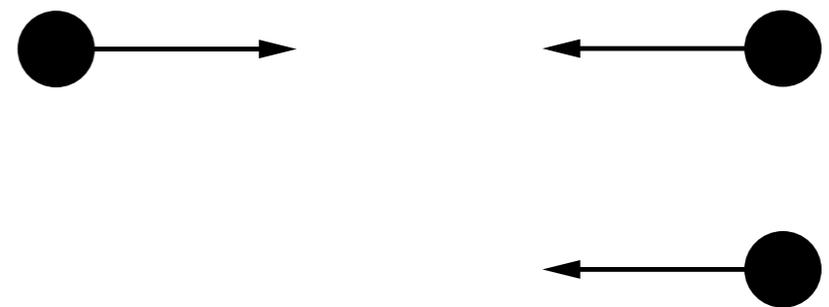
- Self-gravity, irregular geometry : SPH
- High energy explosion : FDM

Artificial viscosity

$$\dot{\mathbf{v}}_i = - \sum_j m_j \left(\frac{P_i}{\rho_i} + \frac{P_j}{\rho_j} + \Pi_{ij} \right) \nabla_i W(\mathbf{r}_i - \mathbf{r}_j, h_{ij})$$

where $\Pi_{ij} = \begin{cases} \frac{-\alpha c_{ij} \mu_{ij} + \beta \mu_{ij}^2}{\rho_{ij}} & \text{if } \mathbf{v}_{ij} \cdot \mathbf{r}_{ij} \leq 0 \\ 0 & \text{elsewhere} \end{cases}$.

- essential for shocks
- turn on in approaching particles
- notorious side effects in a velocity shear (i.e. a keplerian disc)



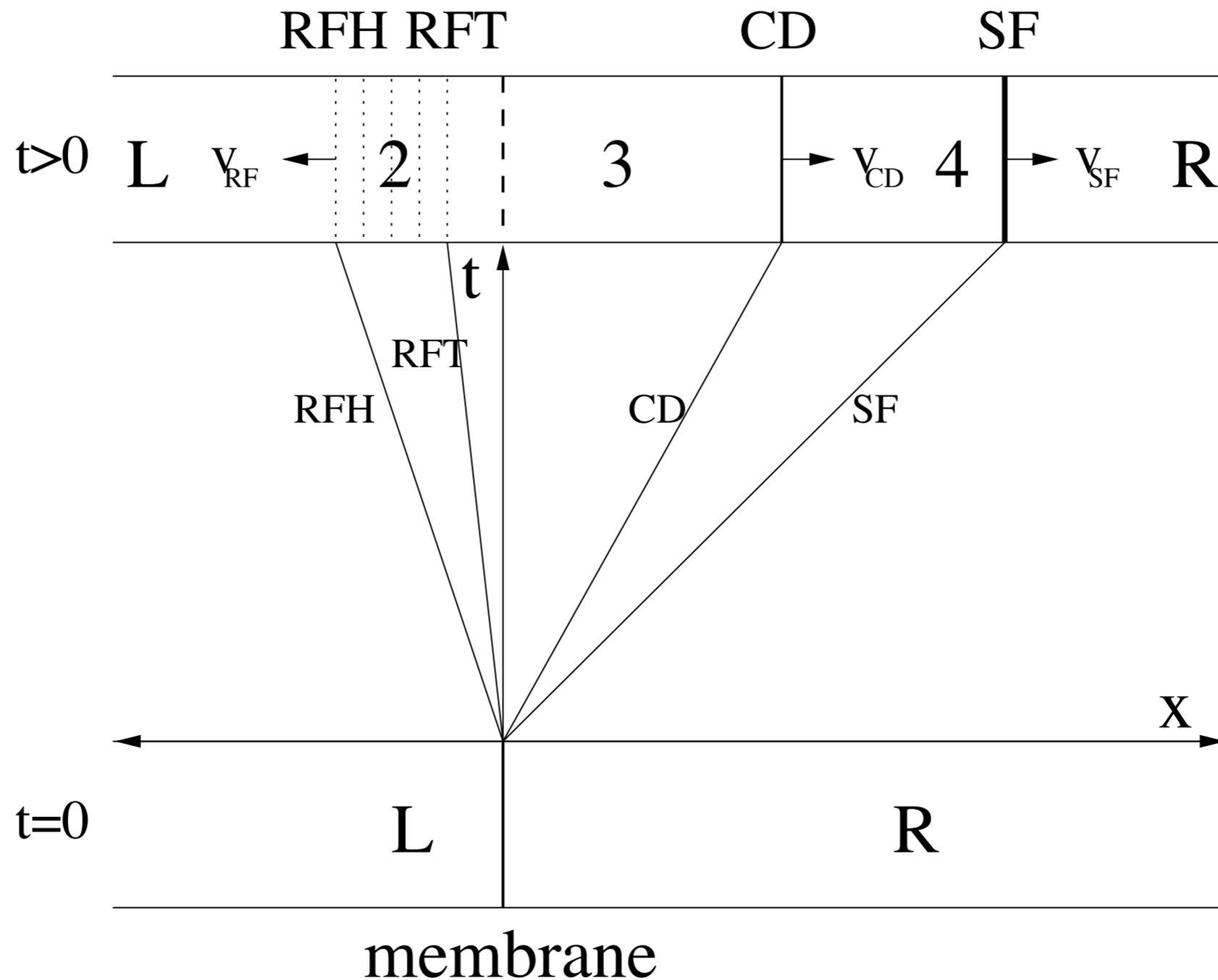
GodunovSPH

Riemann solver instead of the artificial viscosity

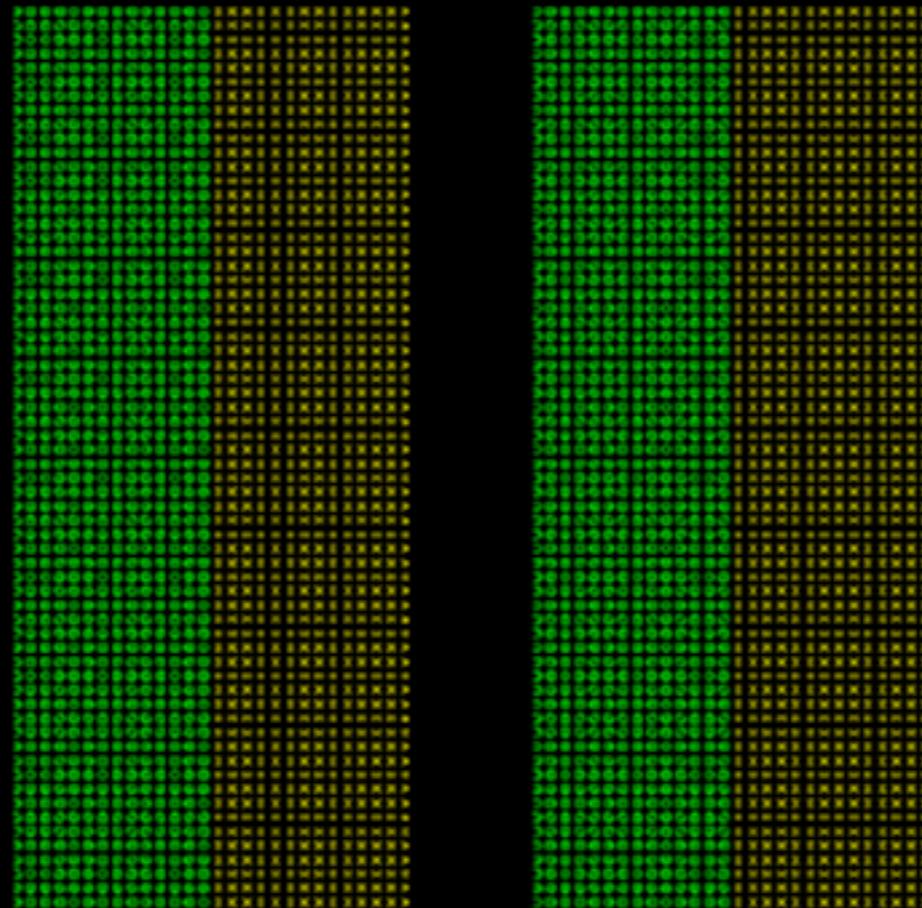
$$\dot{\mathbf{v}}_i = - \sum_j m_j \left(\frac{P_i}{\rho_i} + \frac{P_j}{\rho_j} \right) \nabla_i W(\mathbf{r}_i - \mathbf{r}_j, h_{ij})$$

GSPH can describe (strong) shocks without the AV !

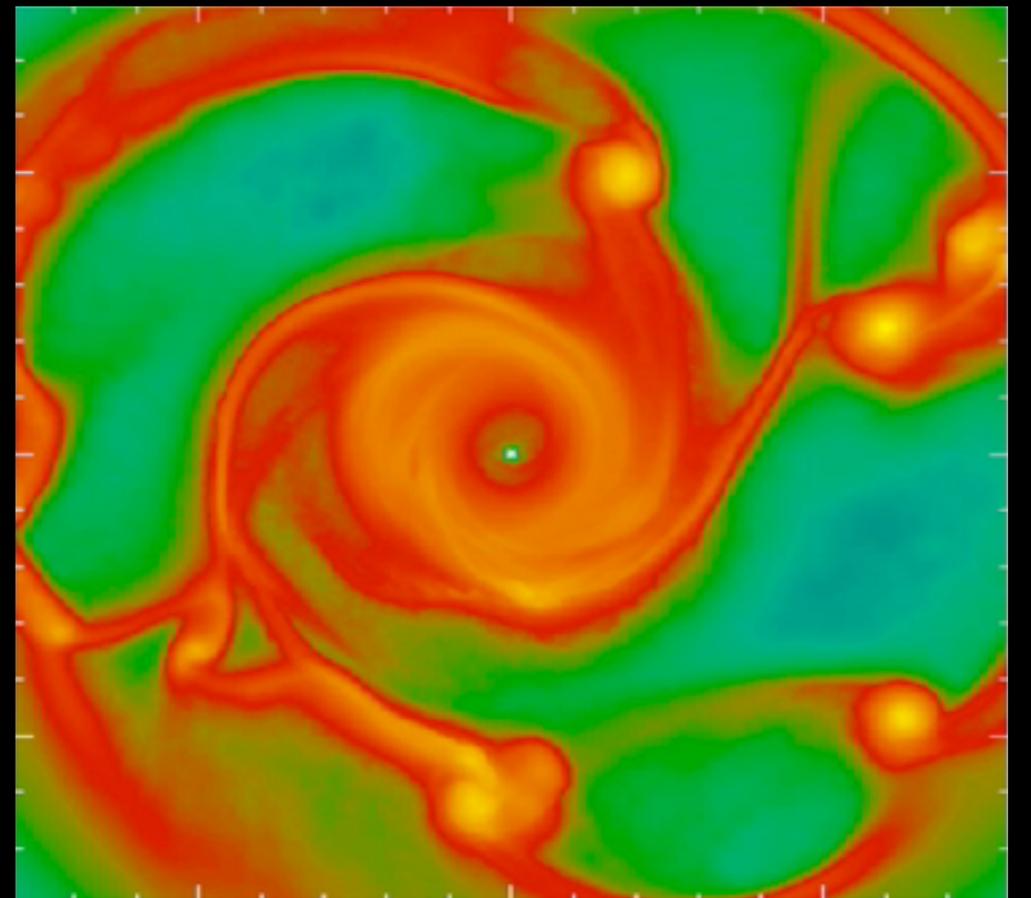
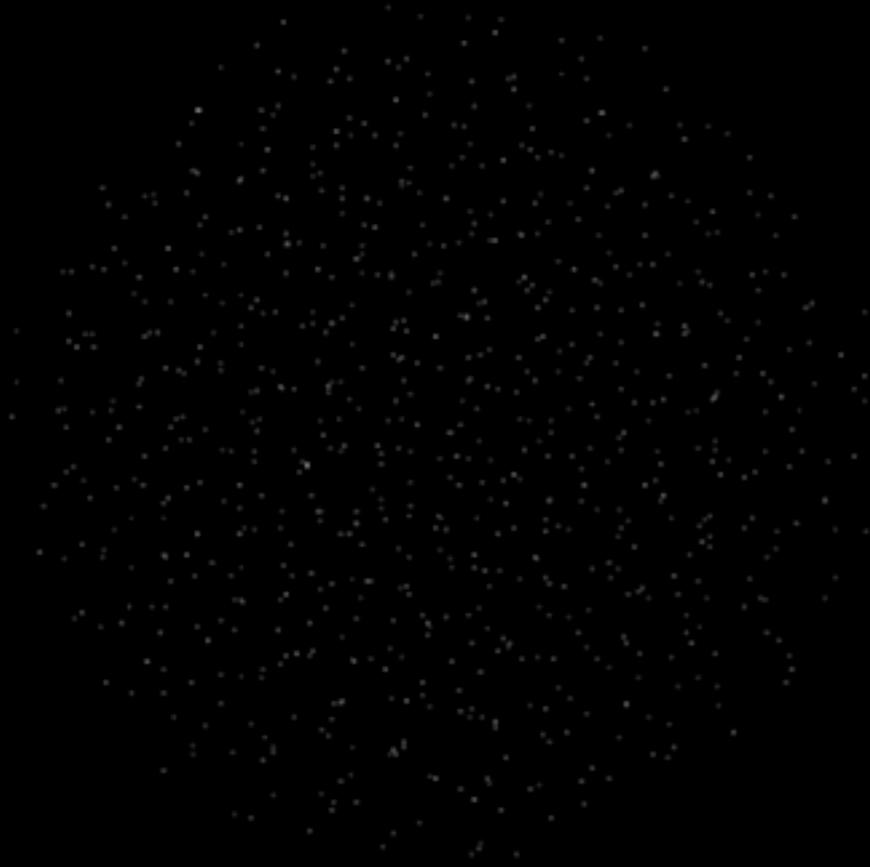
Riemann problem



Side effect of AV



Can we believe these ?



Convergence

Lax-Richtmyer (1956)

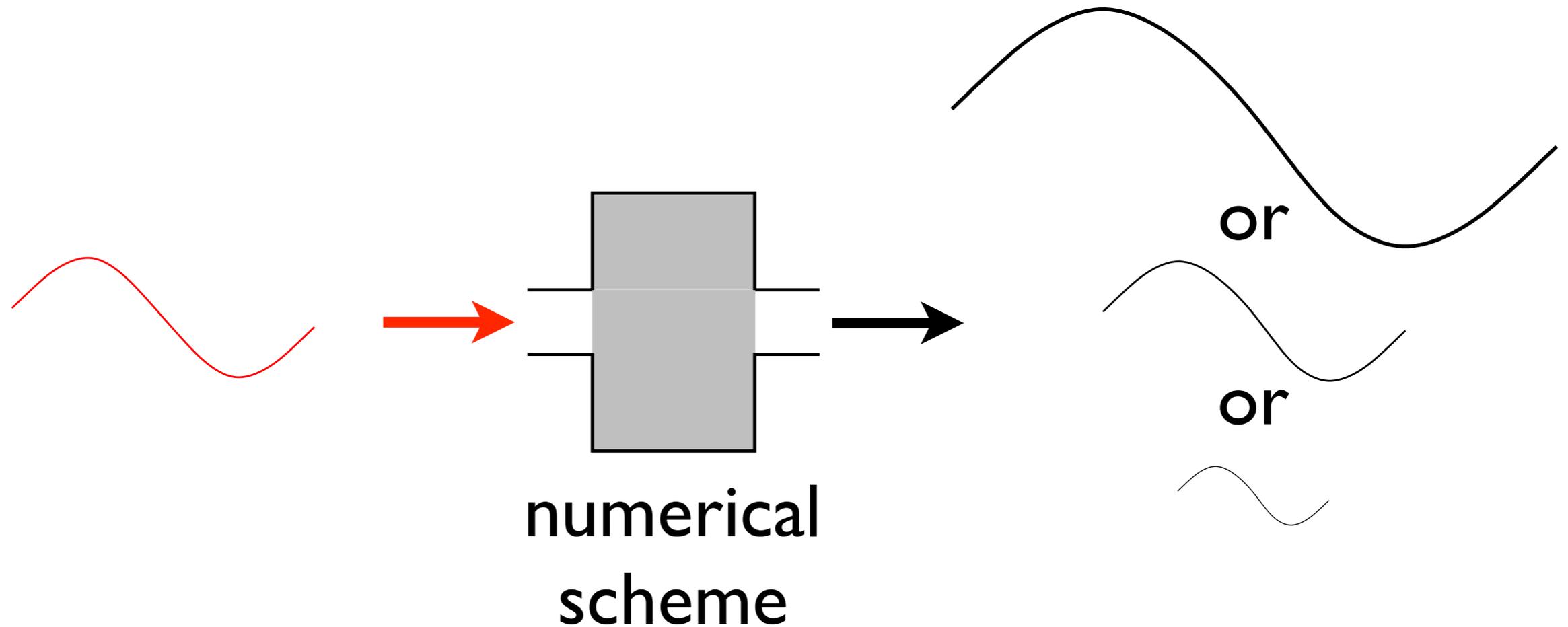
and

Lax-Wendroff (1960)

said that

“**Conservation, Stability and Consistency**
are essential for the **Convergence**”

Conservation



Growth or decay
due to a numerical error



Amplitude Error

$$A_k \equiv A_k(t)$$

GSPH has no amplitude error, but..

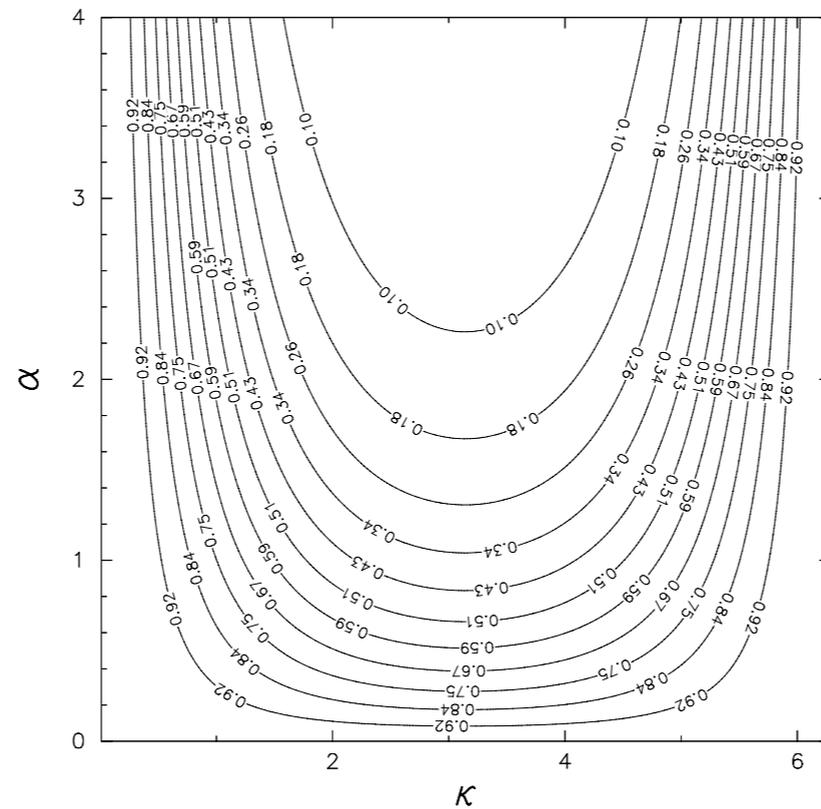
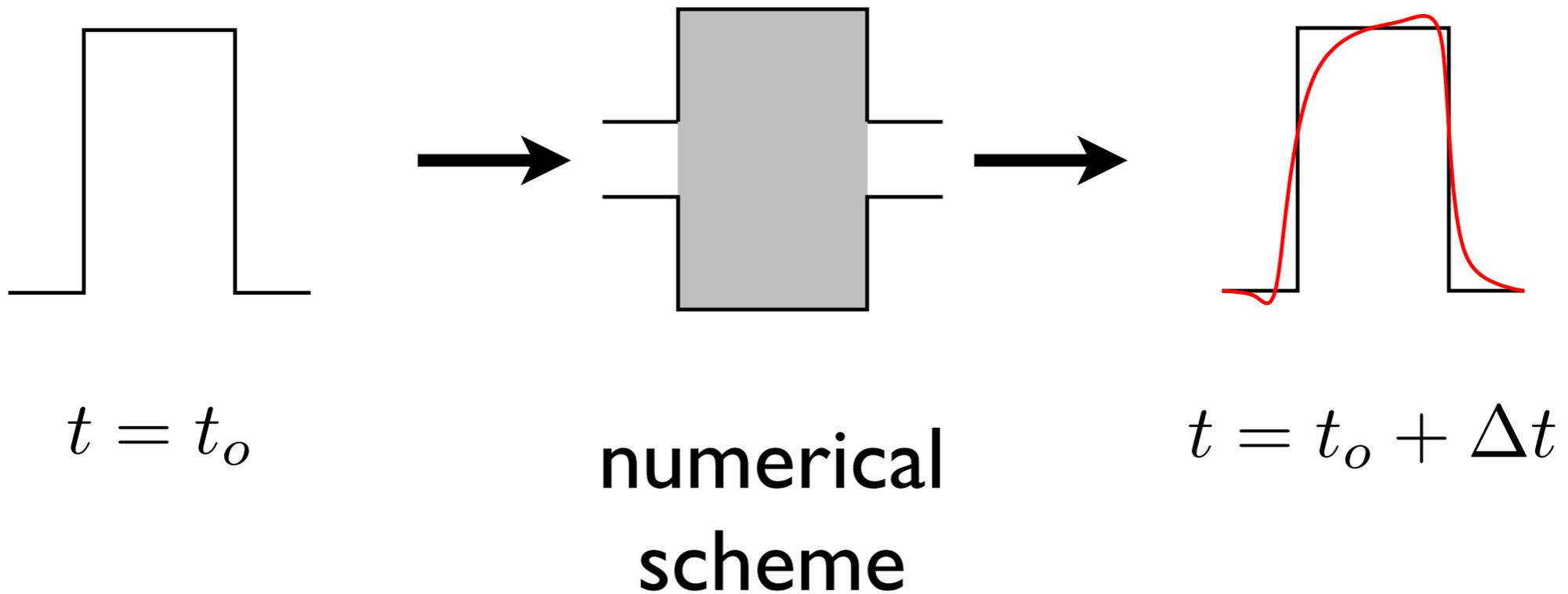


Figure 3.6: Amplitude error contours for isothermal SPH with artificial viscosity on the $\kappa - \alpha$ plane. $b_{SPH,AV}$ is always negative, so the amplitude of the initial perturbation decays with time. The contours show the values at the time, $t = 1$, i.e. just after one period of the wave. The amplitude error depends very weakly on α in the stable region. This means that provided α is set in the stable region, the value of α is not very critical.

Stability

Phase Error, $v \equiv v(k)$



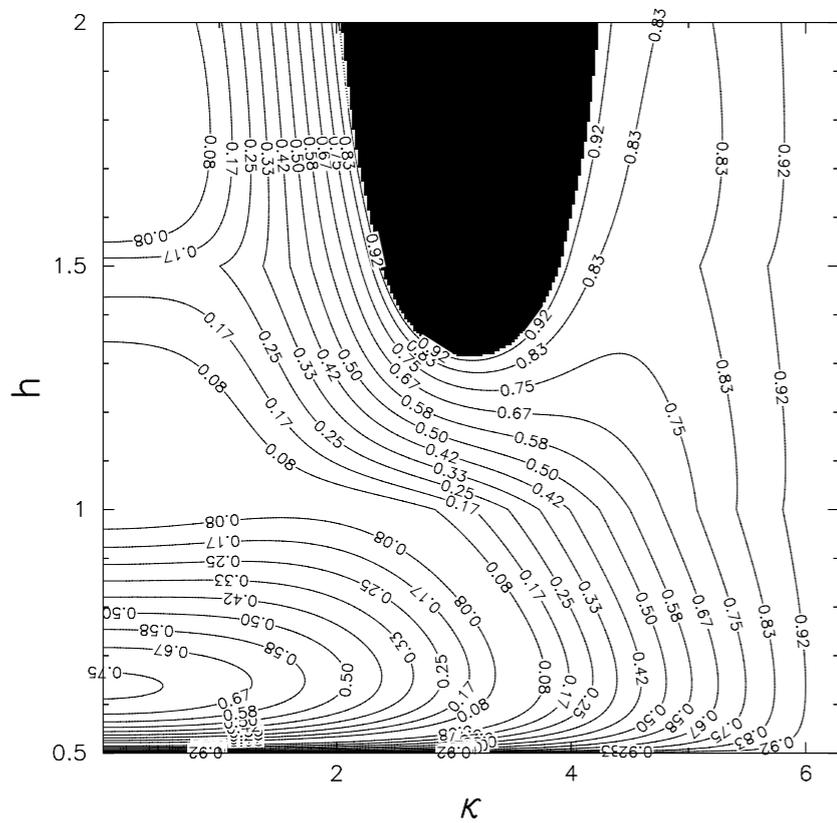


Figure 3.5: Phase error contours for isothermal SPH with artificial viscosity. α is set to 1. The region filled by black in the upper-middle area is the unstable region. In this region, $\omega_{SPH,AV}^2$ is smaller than 0.

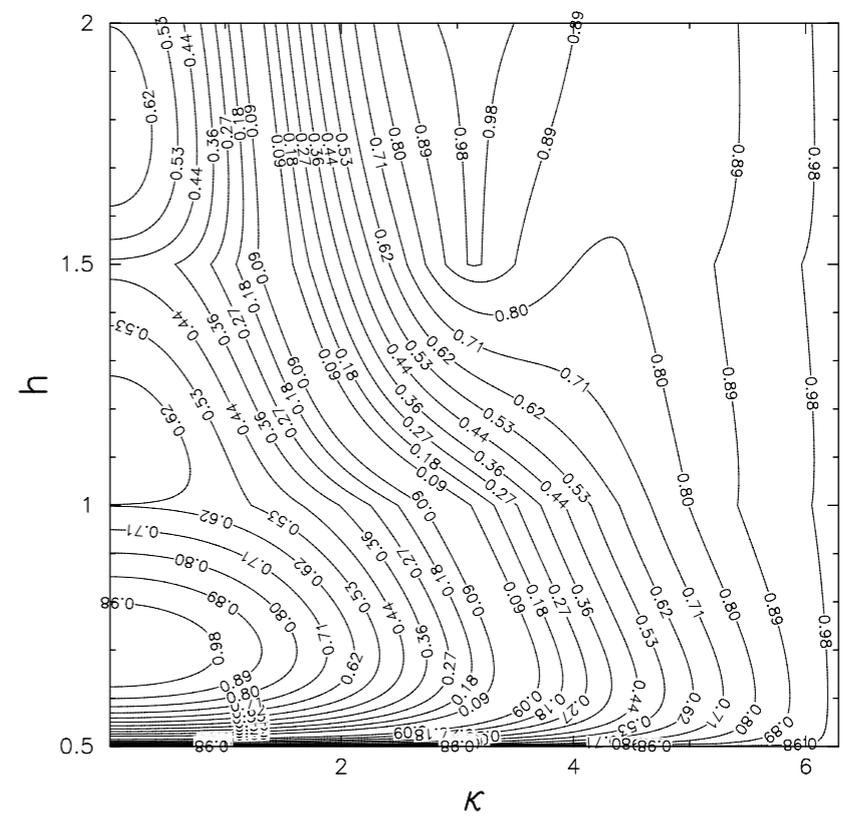
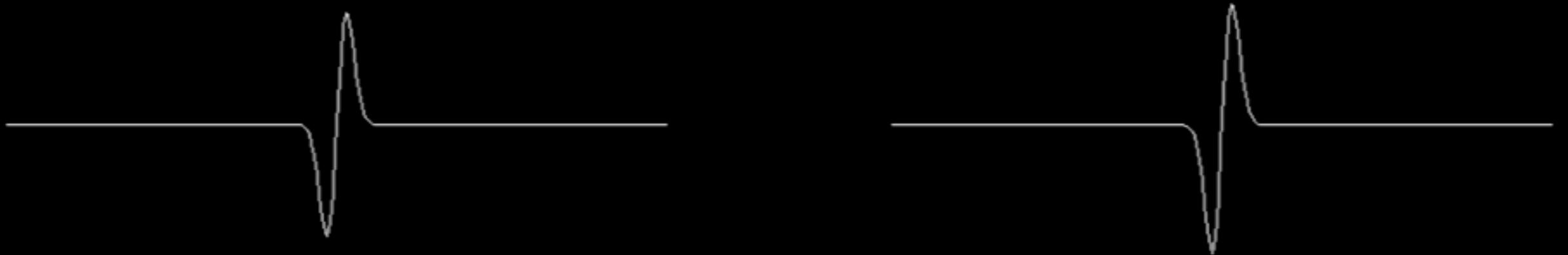


Figure 3.7: Phase error contours for isothermal GPH. At all wavelengths, ω_{GPH}^2 is positive, and the general trend is very similar to the case of SPH without artificial viscosity.

Sound wave



Phase error-free scheme is important in inkjet printers

$$q_i^{n+1} = q_i^n - \frac{u}{4} \frac{\Delta t}{\Delta x} (q_{i+1}^n + 3q_i^n - 5q_{i-1}^n + q_{i-2}^n) - \frac{u^2}{4} \frac{\Delta t^2}{\Delta x^2} (q_{i+1}^n - q_i^n - q_{i-1}^n + q_{i-2}^n)$$

Fromm's scheme (1984)

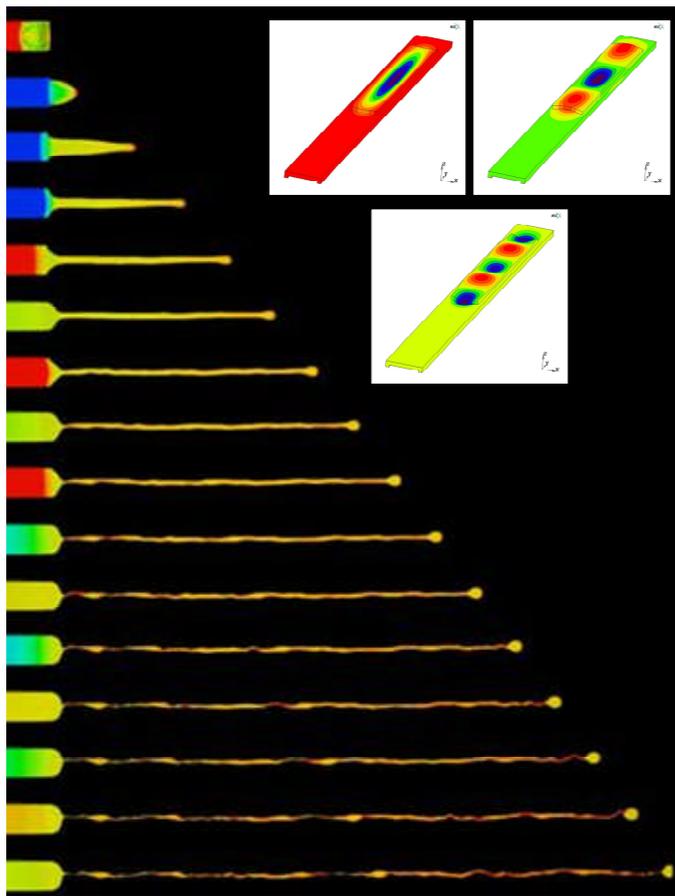


Figure 2. Inter-disciplinary simulation of piezoelectric inkjet formation. Time interval is 6% of firing cycle.

Zeng (2009)

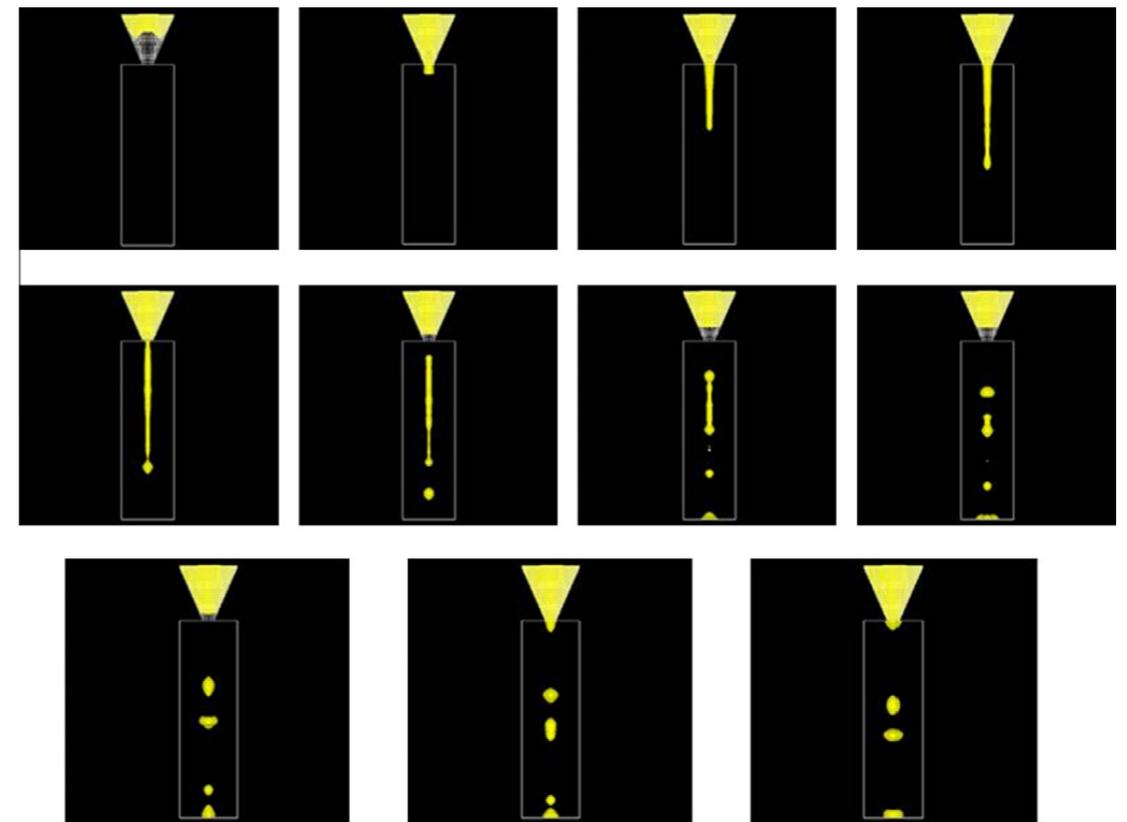
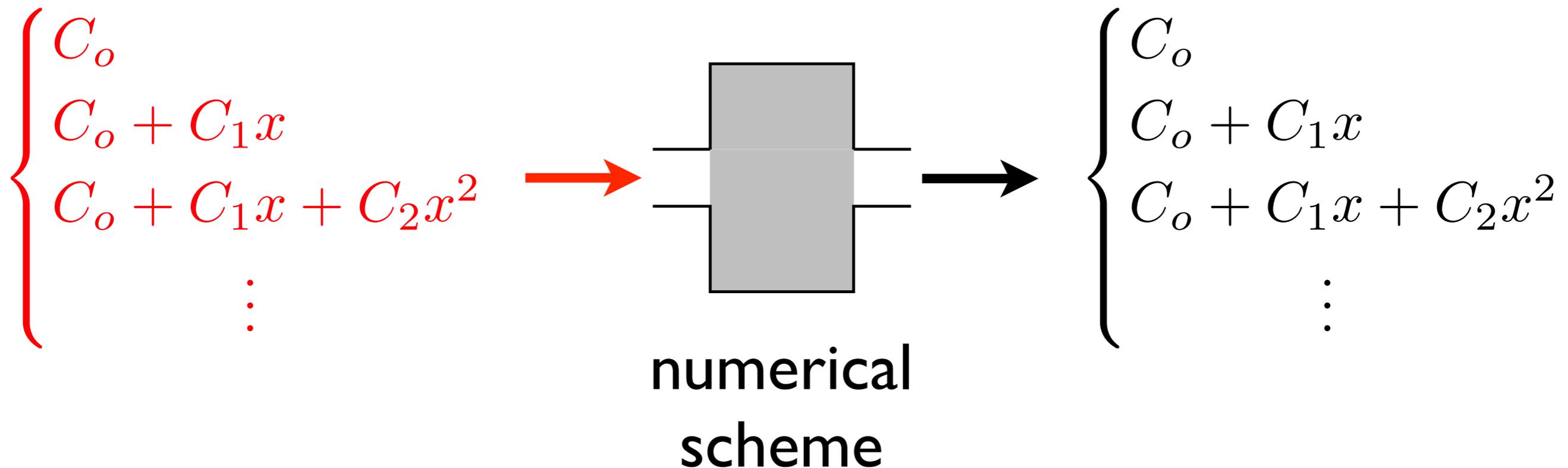


Fig. 9. Numerical results of column droplet.

Chang (2011)

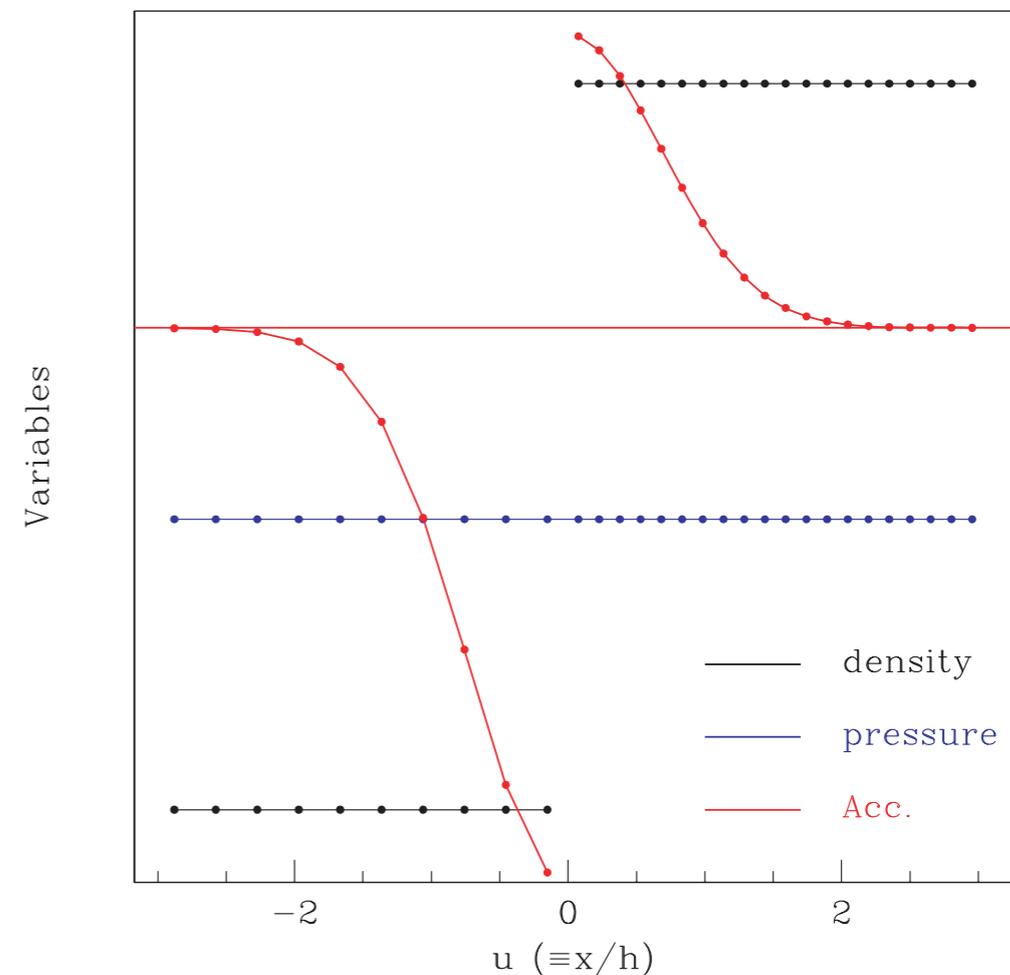
Consistency

- Resemblance of PDE and the numerical equations
- Analysis of the truncation error
- Order of accuracy



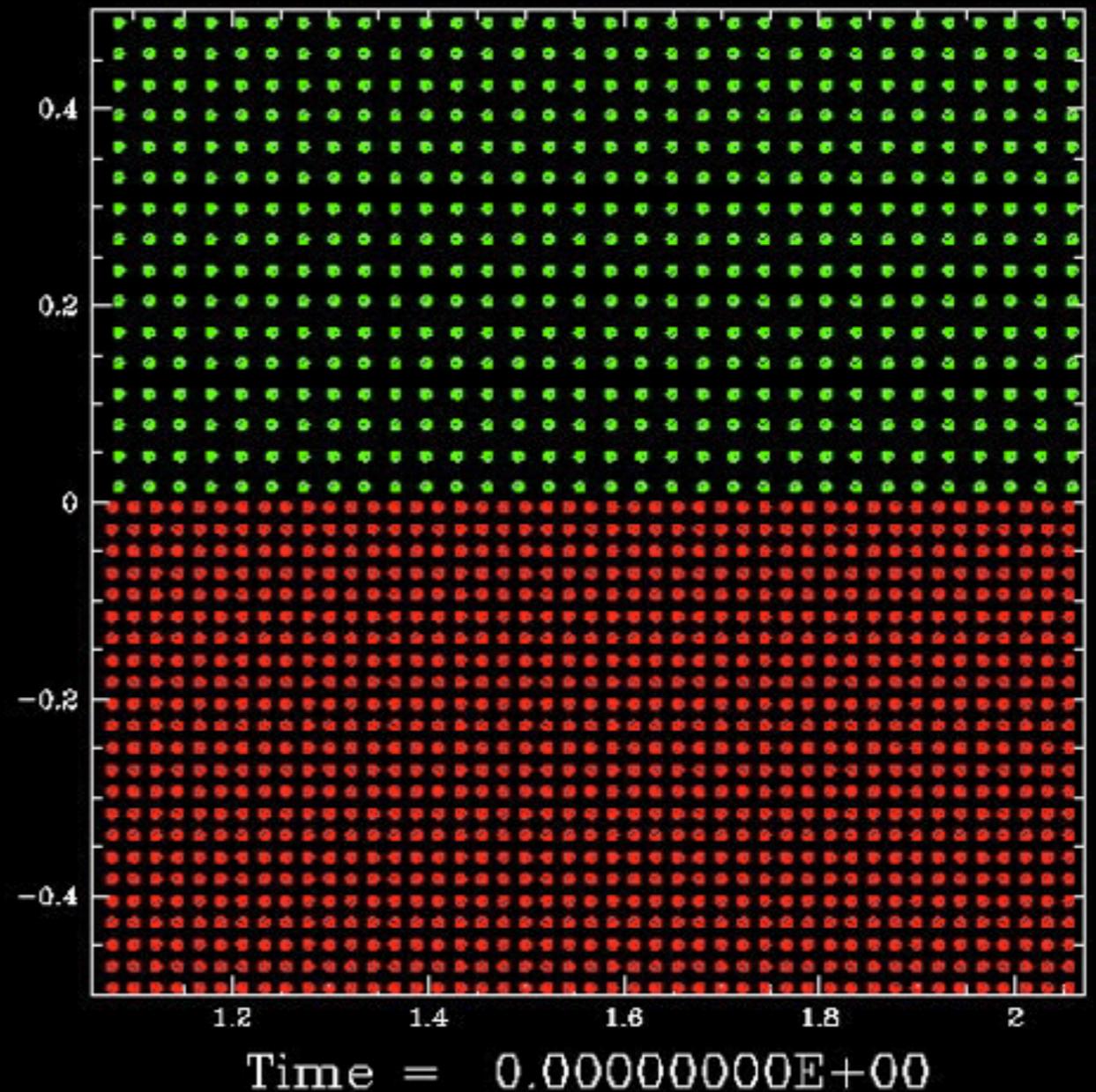
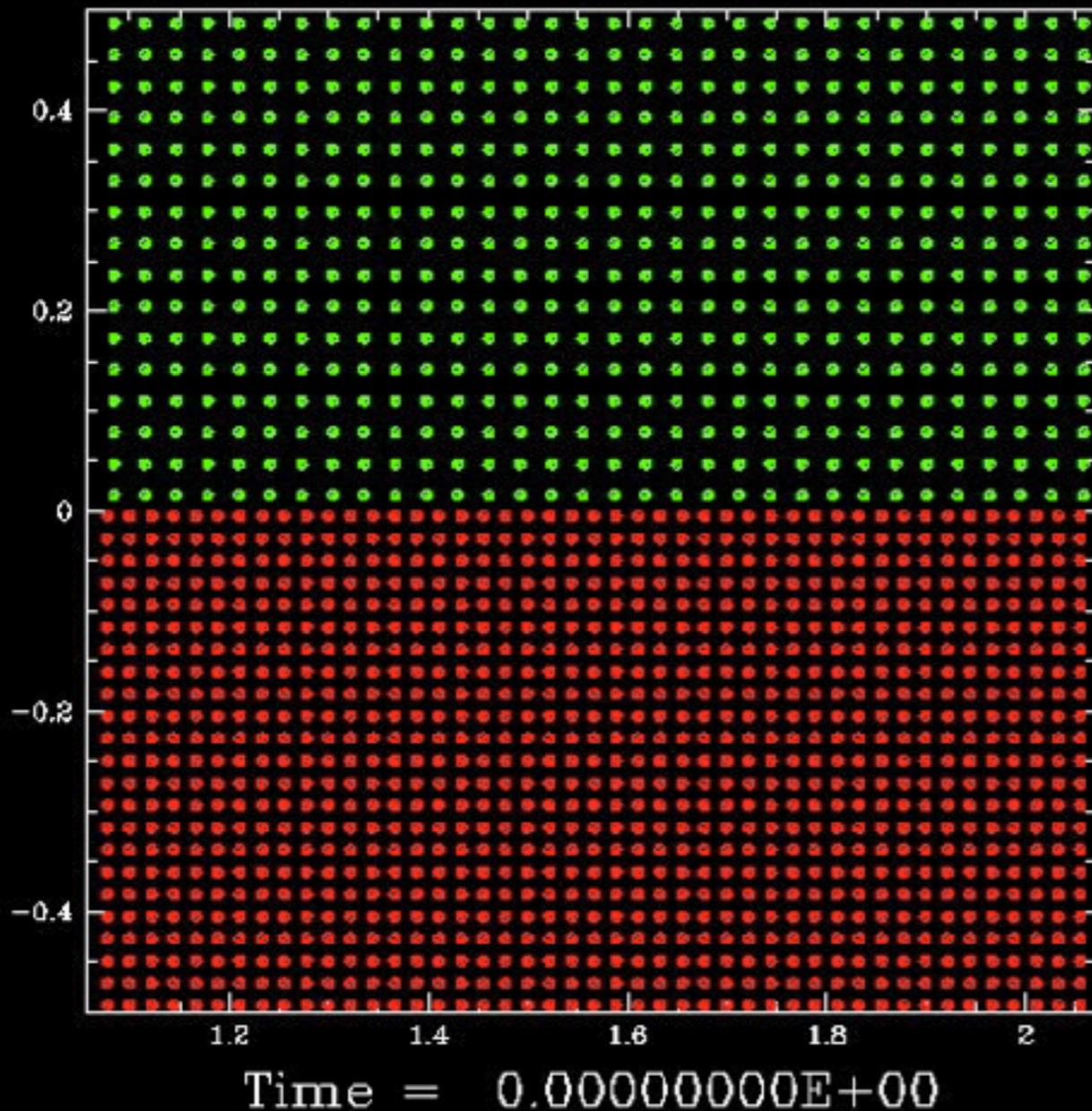
Numerical Surface tension

- SPH is inconsistent.
- The numerical surface tension appears due to the inconsistency.



Cha et al. 2010

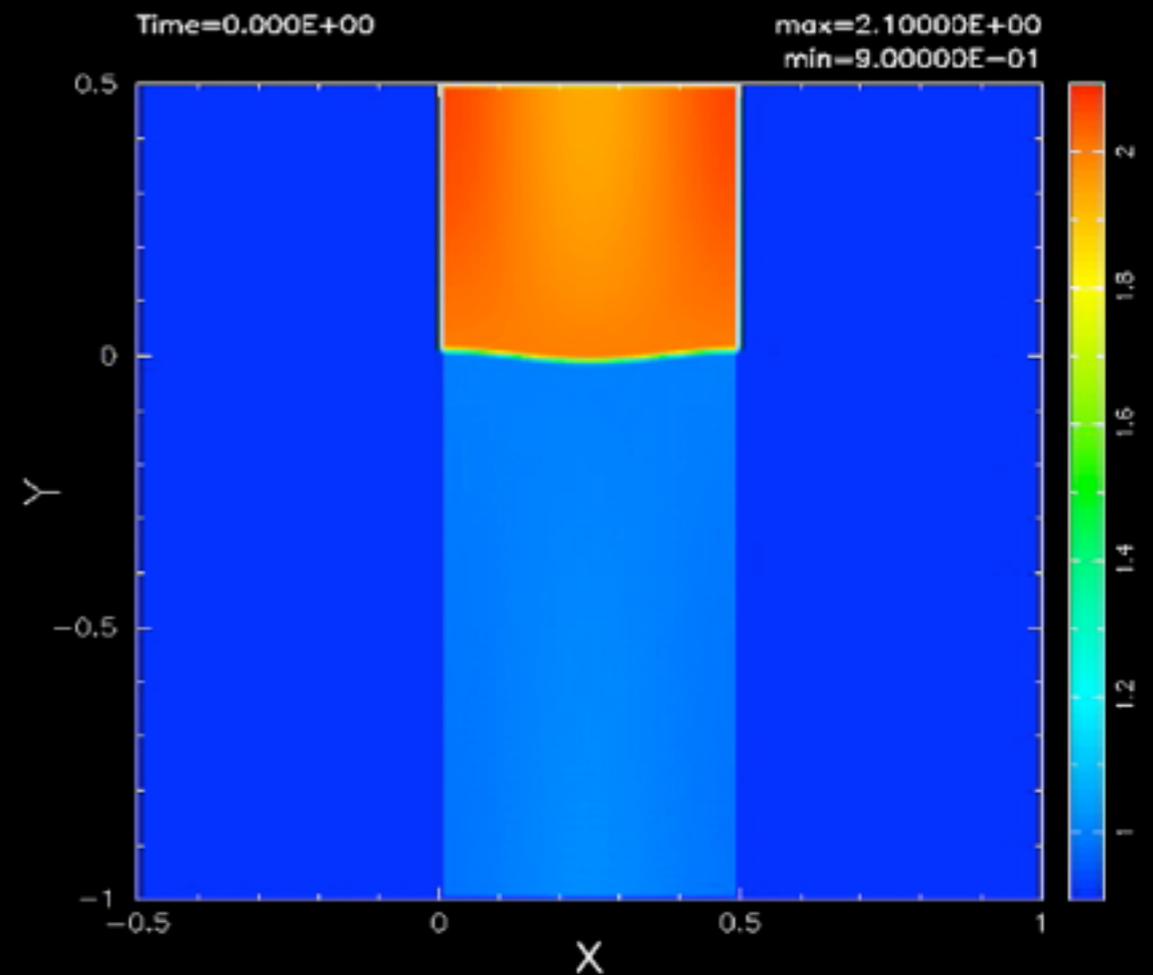
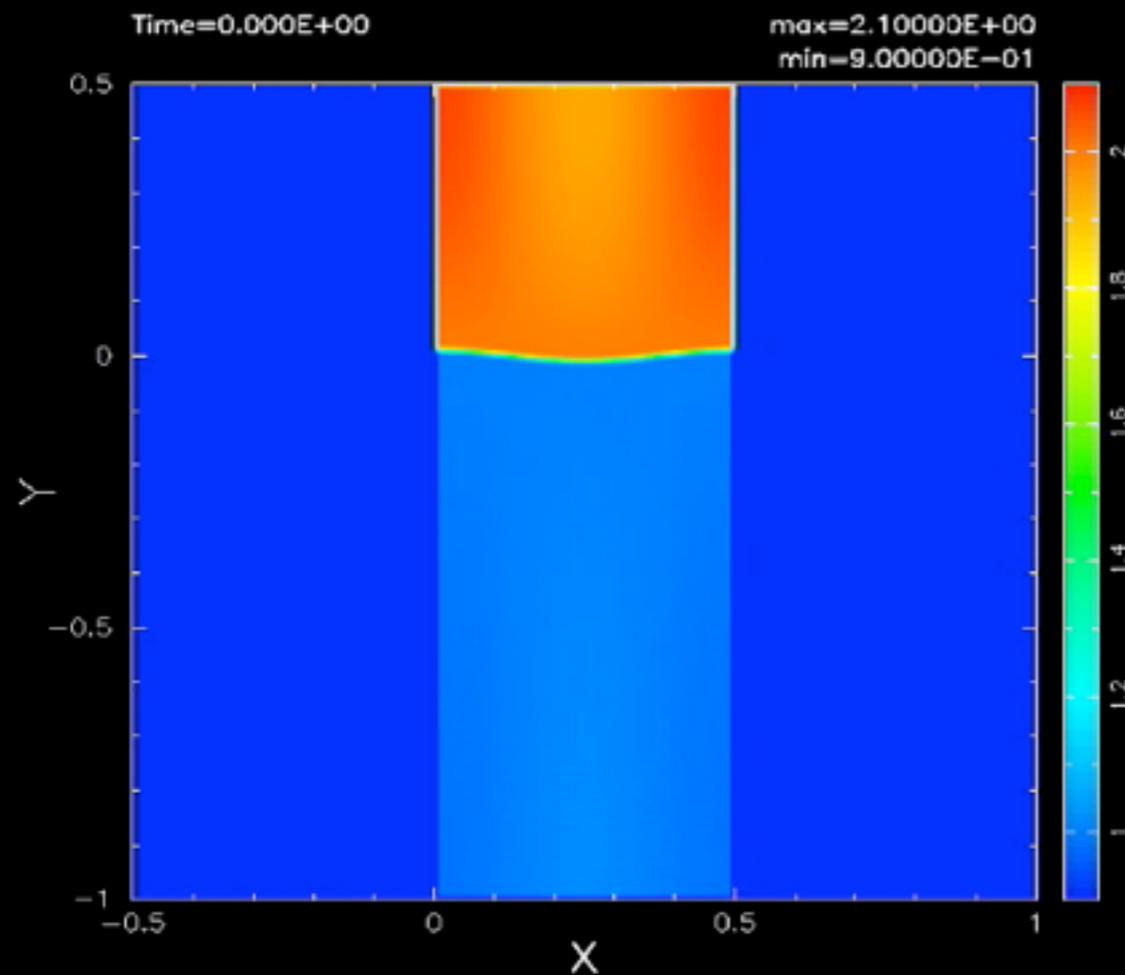
Numerical surface tension



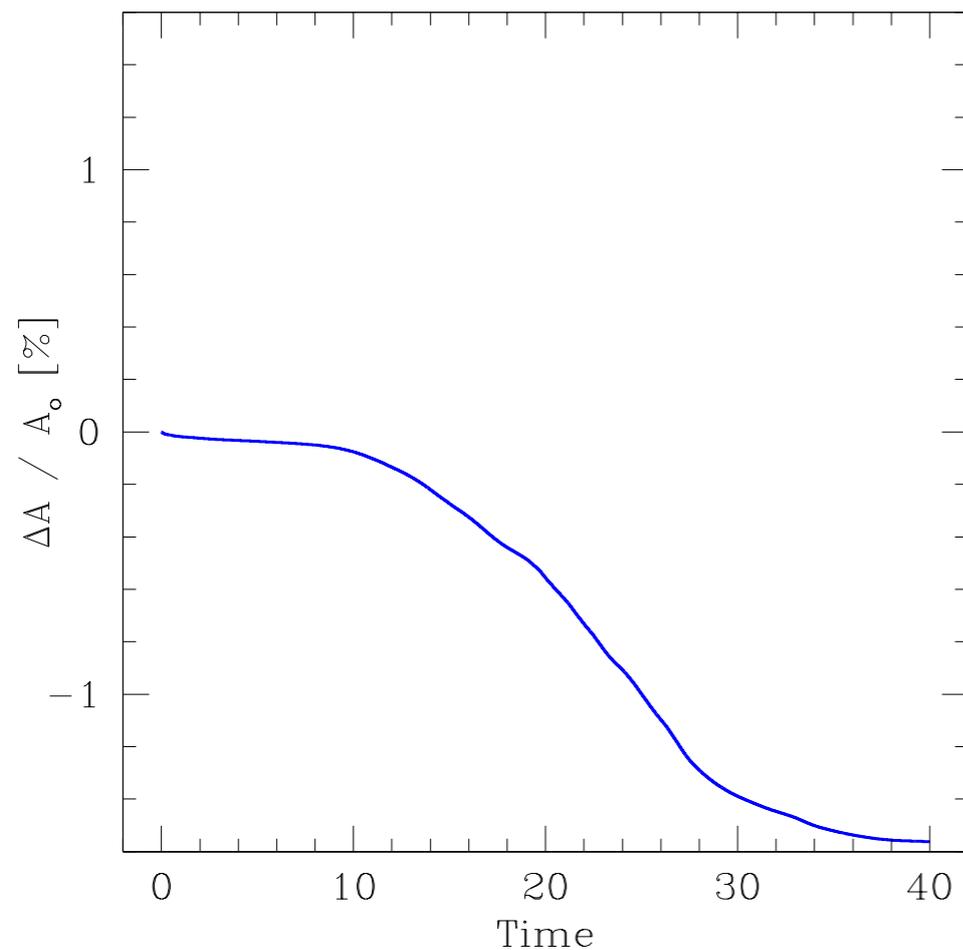
GSPH .vs. SPH

| | GSPH | SPH |
|--------------|---------------|----------------|
| Conservation | ○ | × |
| Stability | ○ | × |
| Consistency | First-ordered | Zeroth-Ordered |

Rayleigh-Taylor Inst.

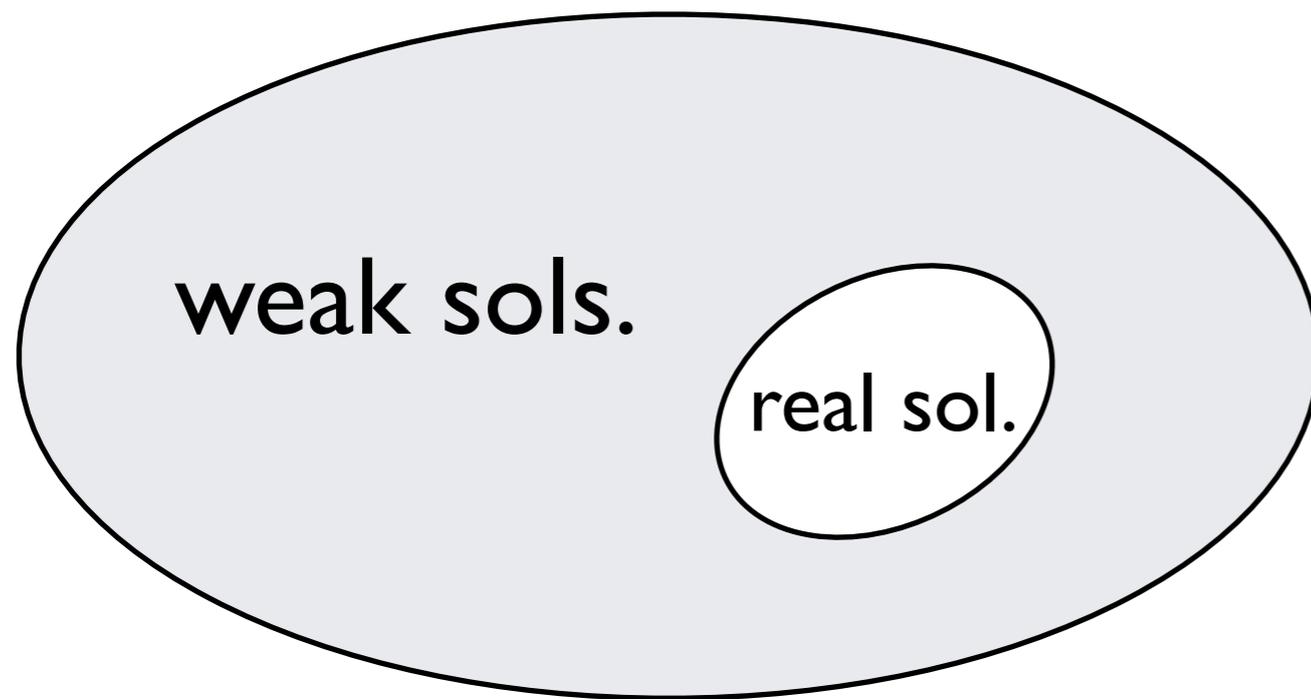


What was wrong ?



- Lax-Richtmyer and Lax-Wendroff said **necessary** conditions.
- Convergence to a **weak solution !**

Weak solutions



Entropy condition

(Olenik, 1963)

$$\frac{ds}{dt} \geq 0$$

Thermal compatibility

the entropy of a fluid, $S = c_v \ln \left(\frac{P}{\rho^\gamma} \right)$

$$\frac{Du}{Dt} = -\frac{P}{\rho} \nabla \cdot \mathbf{v}$$

$$\frac{D\rho}{Dt} = -\rho \nabla \cdot \mathbf{v}$$

$$\therefore \frac{Du}{Dt} = -P \frac{DV}{Dt}$$

Density estimation

$$\rho_i = \sum_j \frac{m_j}{V_i} \int \frac{1}{\rho(\mathbf{x})} W((\mathbf{x} - \mathbf{x}_i), h_i) W((\mathbf{x} - \mathbf{x}_j), h_j) d\mathbf{x}$$

tcGSPH

Assume $\frac{1}{V_i} \simeq \rho(\mathbf{x}) \simeq \rho_i$, then

GSPH

$$\rho_i = \sum_j m_j \int W((\mathbf{x} - \mathbf{x}_i), h_i) W((\mathbf{x} - \mathbf{x}_j), h_j) d\mathbf{x}$$

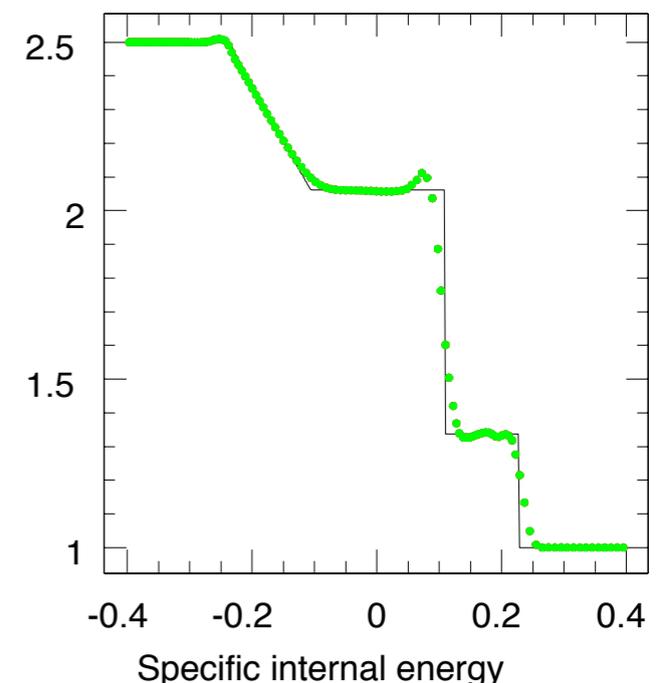
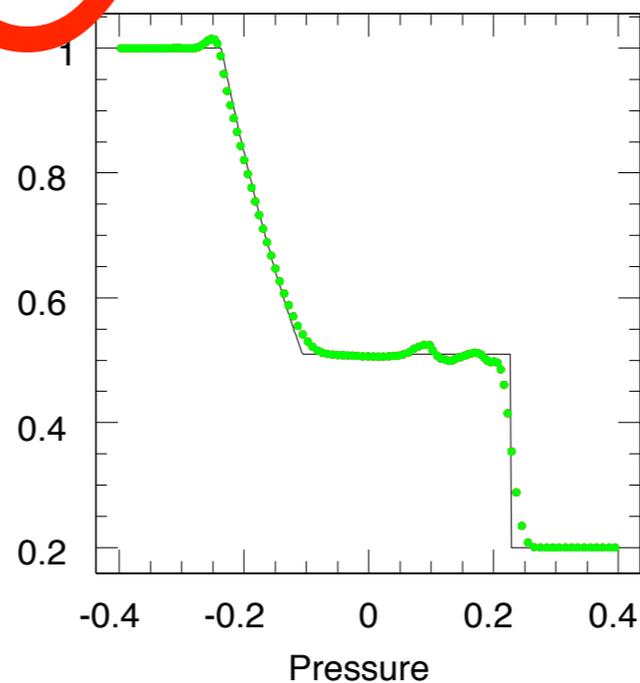
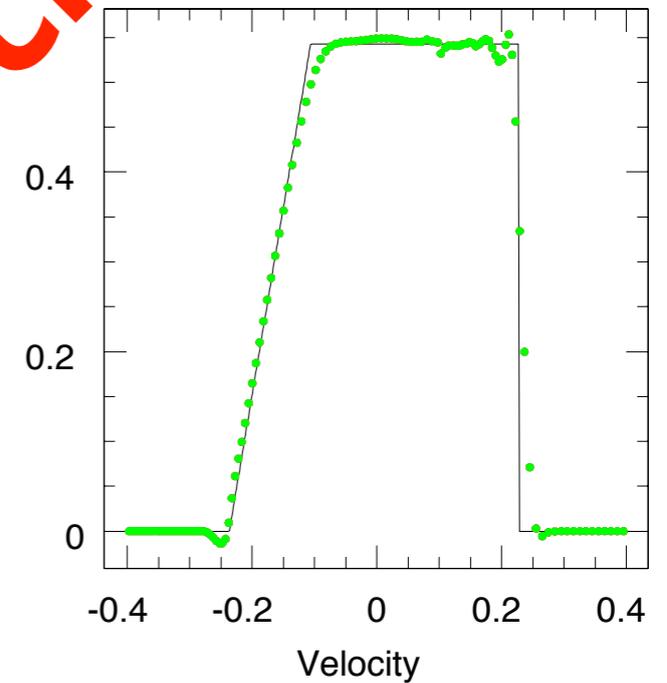
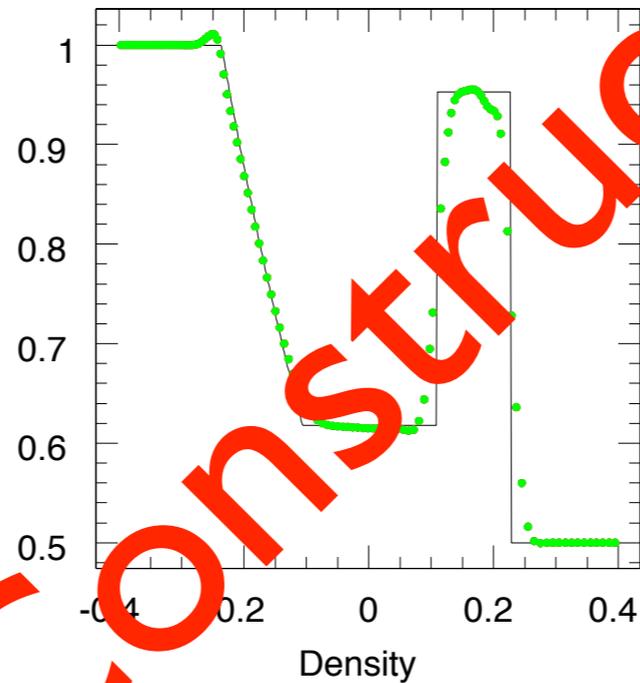
Assume $\frac{1}{V_i} \simeq \rho(\mathbf{x}) \simeq \rho_i \simeq \rho_j$, then

SPH

$$\rho_i = \sum_j m_j W(\mathbf{x}_i - \mathbf{x}_j, h_{ij})$$

Shock tube test

- The interactions between the two different fluid.
- It is called Sod test (Sod 1978)
- A good test problem



SPH, GSPH & tcGSPH

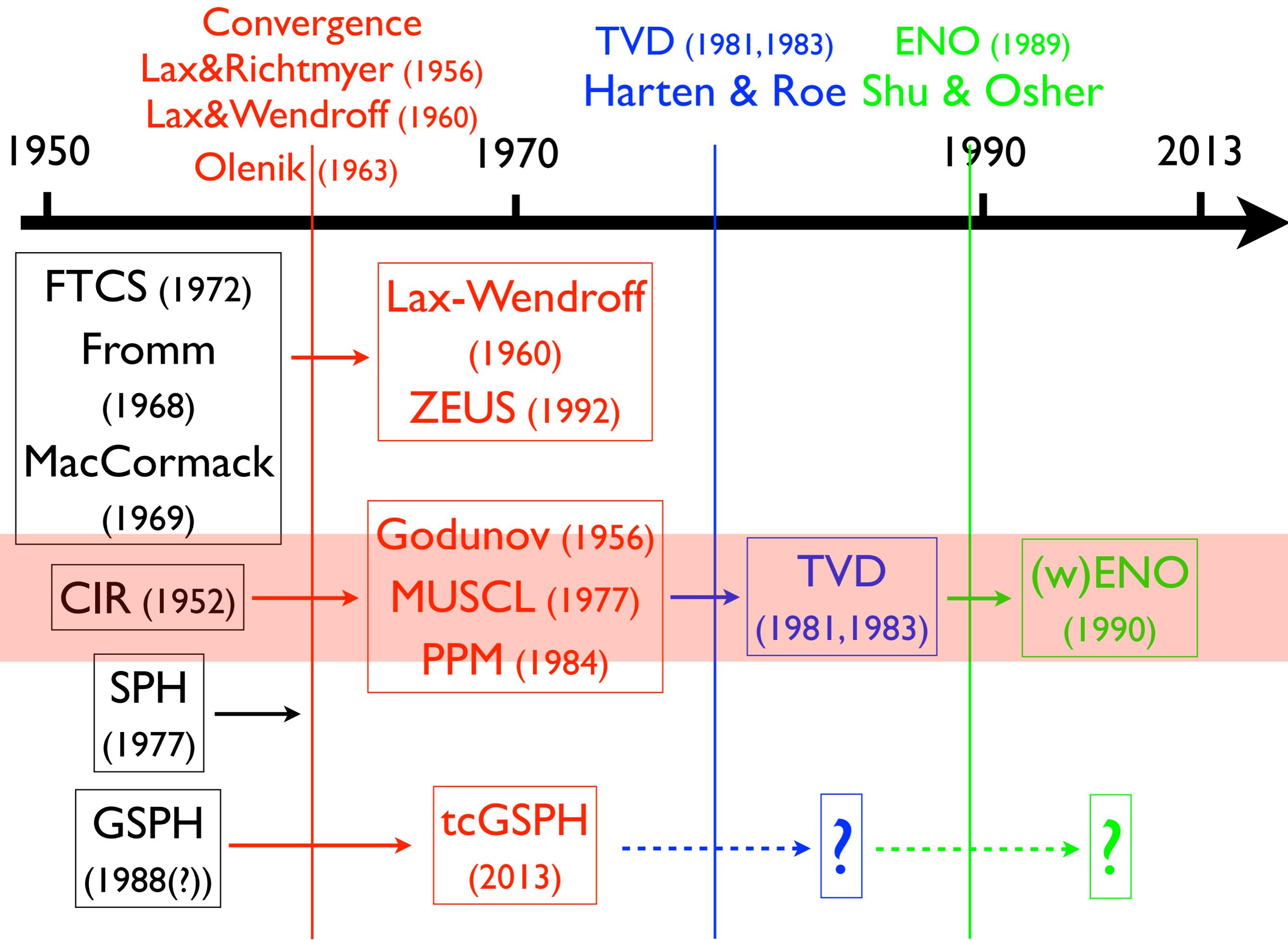
| | SPH | GSPH | tcGSPH |
|-------------------|----------------|---------------|---------------|
| Conservation | X | ○ | ○ |
| Stability | X | ○ | ○ |
| Consistency | Zeroth-Ordered | First-ordered | First-ordered |
| Entropy condition | X | X | ○ |

Why are the results different ?

They didn't obey the rules.

Can I believe my simulation ?

Yes I can.



Future

- Nonlinear, nonisothermal stability analysis
- Total Variation Diminish (TVD)
- Non-oscillatory property
- Different parallelization (i.e. GPU)

tcGSPH is

- working.
- better.

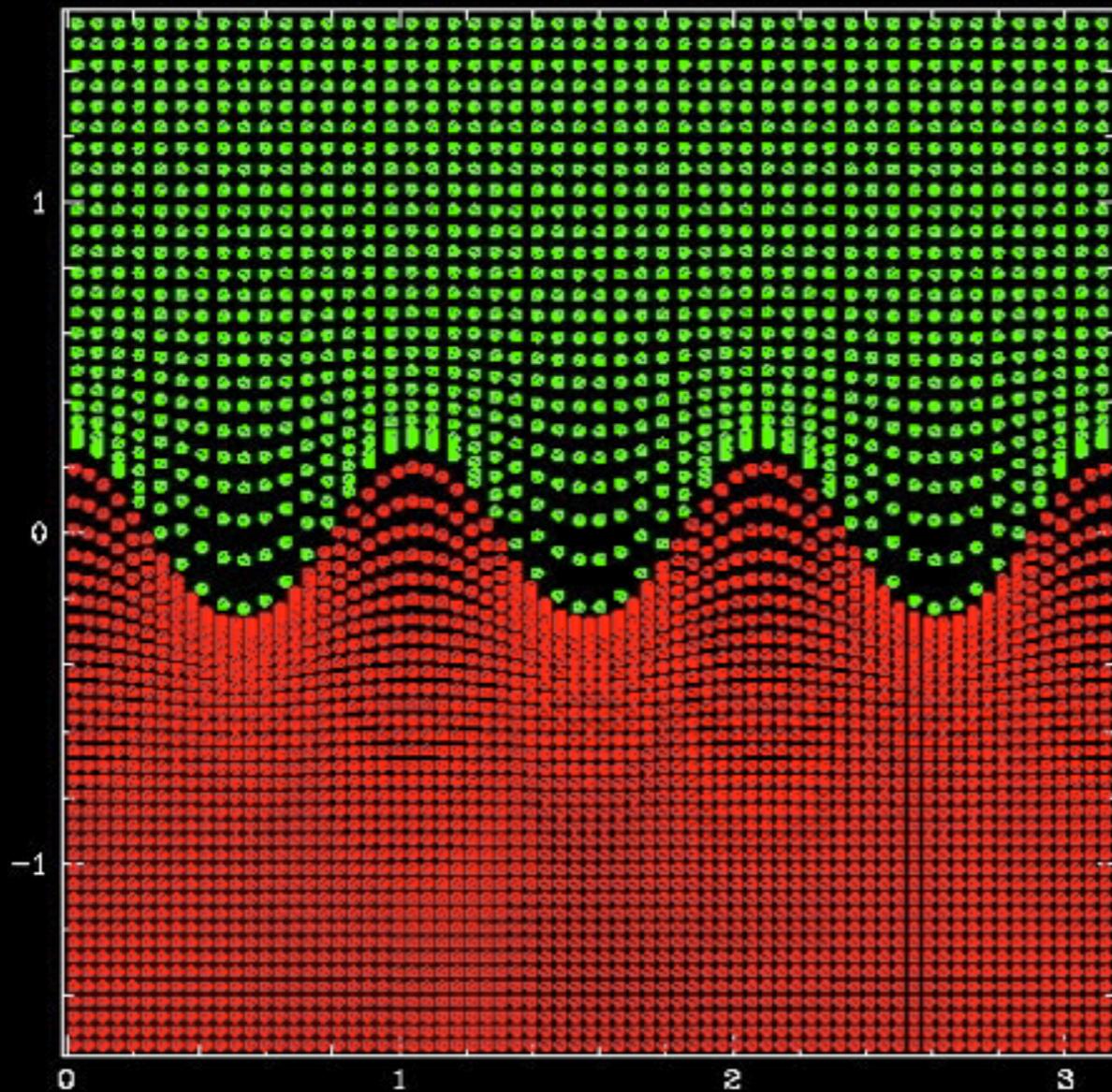
Public code for CVs

- Convergence scheme : high resolution scheme
- Radiation HD by the diffusion approximation

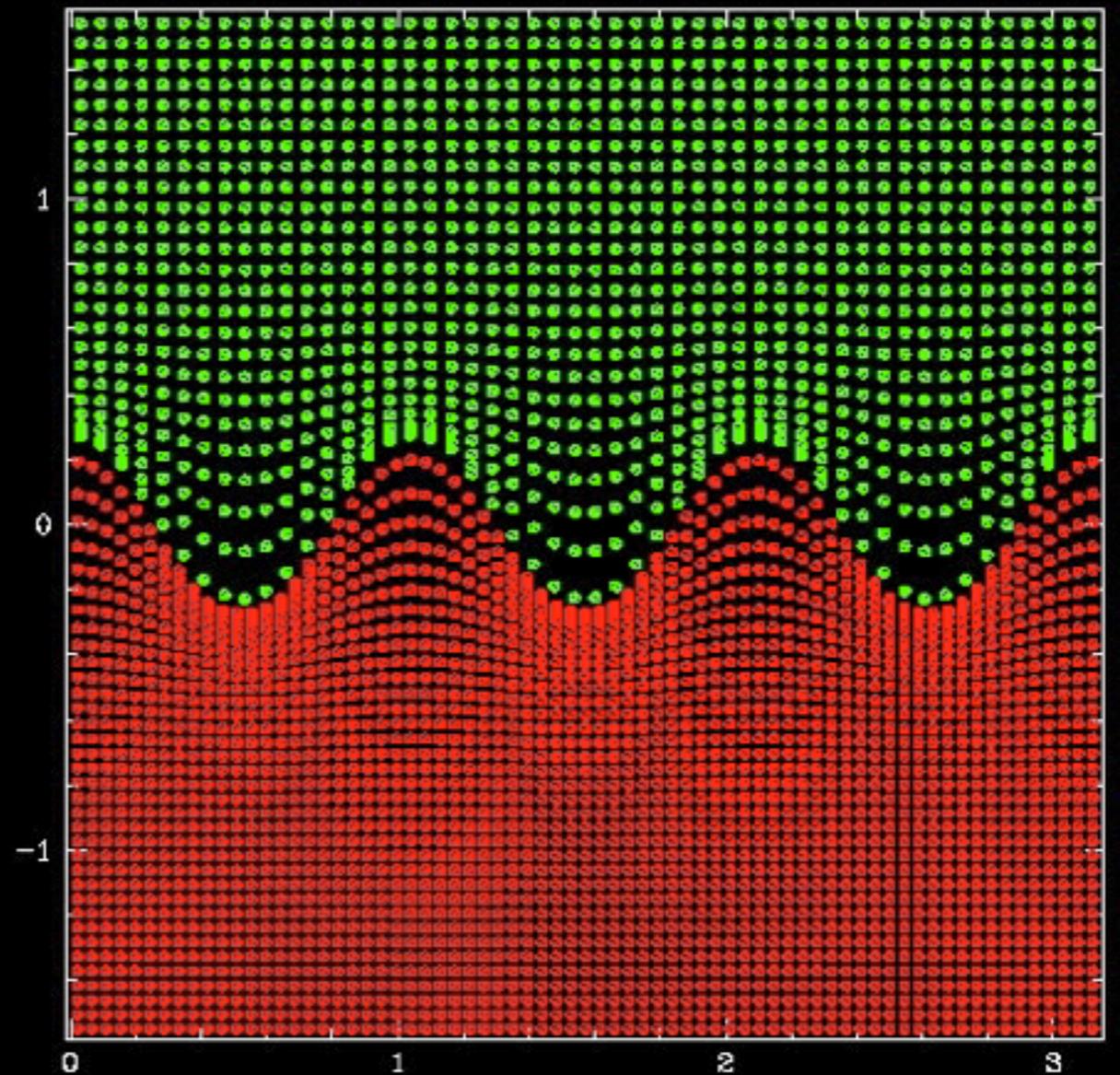
Any participation should be welcomed.

- Fully 3-dimensional self-gravity code
- Parallelization with MPI2-libraries
- Magneto hydrodynamics (optional)
- Documentation : HTML, (La)TeX, UNIX man...
- Scientific visualization

Solenoidal field



Time = 0.00000000E+00



Time = 0.00000000E+00

